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Image Nat Astron 4, 913–915 (2020)

Motivation

UHE neutrinos are a key science target

- Messengers for UHE particle accelerators
- Radio detection activities for neutrinos going strong



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Acoustic neutrino detection an "old dream"

- Askaryan effect creates radio and acoustic emission
- First studies reach back to DUMAND times
- Technology and background were major obstacles



Experimental landscape

Experiment	Location	Medium	Sensor	Host
			Channels	Experiment
SPATS [37, 38]	South Pole	Ice	80	IceCube
Lake Baikal [39]	Lake Baikal	Fresh Water	4	Baikal Neutrino
				Telescope
OVDE 40	Mediterranean Sea (Sicily)	Sea Water	4	NEMO
AMADEUS 41	Mediterranean Sea (Toulon)	Sea Water	36	ANTARES
ACoRNE 42	North Sea (Scotland)	Sea Water	8	Rona military array
SAUND 43	Tongue of the Ocean	Sea Water	7/49(*)	AUTEC military
	(Bahamas)			array
SMO	Mediterranean Sea	Sea Water	10	NEMO Phase-II
sensor array [44]	(Sicily)			prototype
KM3NeT	Mediterranean Sea	Sea Water	under	KM3NeT
sensor array [45]	(Sicily, Toulon)		constr.	

Reference <u>http://dx.doi.org/10.1051/epjconf/201921601001</u> (2019)

(*) The number of hydrophones was increased from 7 in SAUND-I to 49 in SAUND-II

NEW

ANDIAMO Adriatic Sea

Sea Water proposed -

Marinelli et al (2022)

SFL

Experimental landscape

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Is it wor	th to re-inves	tigate thi	s effor	t?	
• Mo • Key	dern computi /: Sensitivity ii	ng, mach n the sub-	ine leai 100 Pe	rning V regime	

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The study of acoustic neutrinos requires simulation

- I: Shower development, energy deposition
- II: Velocity potential, medium effects
- III: Acoustic pulse
- Future: Sensor arrays, event reconstruction, sensitivities

CORSIKA 8 simulation framework [1][2]

- Encouraging support, had running software after a few hours
- Shower development possible in both water, ice
- Started new branch for acoustic shower developments



Image Bevan et al., ACORNE Coll., NIM:A (2009)

[1] Huege et al., *PoS(ICRC2023)* 310[2] Engel et al., *Comput Softw Big Sci* 3, 2 (2019)

Simulation



Standard CORSIKA8 shower outputs

Simulation



Standard CORSIKA8 shower outputs $(u_{\mu} \text{ with } 10^8 \text{ GeV, NC})$

			Goal	T.5 Conega + E 7.5 ×10 μ 5.0 Sigma- μ 2.5 1 μ 2.5 1 μ	K(1)0 Eambda Sigma + K(2)0 Ki Ki Ki D2 T3	
	Acous	stic n	eutring	o pulse	S	

Step 1: Radial shower profiles

- Writing radial energy deposits (with cut particles excluded)
- **Goal:** Radial, longitudinal energy density distribution



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Step 2: MC shower energy density

- Sample radial energy density to arrive at 3D energy density
- Produce MC shower sampled from that density



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Step 3: Velocity potential, acoustic pulse [3]

- Sample propagation times from MC shower cells to sensor position in medium
- Velocity potential: Scale with total energy, medium properties





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- Velocity potential: Scale with total energy, medium properties
- Acoustic pulse: Differentiate



Simulation



No seawater correction applied.

SFU

Simulation



No seawater correction applied.

Summary

Successful generation of acoustic neutrino signatures in water based on CORSIKA 8 shower simulations

- More effort needed to verify results against existing work
- Systematic study of acoustic signatures (~ energy, observation angle) needed
- Intention to eventually develop an acoustic module for CORSIKA 8

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Future work will target next steps for acoustic neutrino arrays

- Verification of acoustic pulses with literature
- Development of an acoustic neutrino simulation toolkit
- Scalable acoustic sensor technology R&D
- Event reconstruction, sensitivity studies
- Joint detection capabilities with optical, radio

