

# Study of acoustic neutrino detection in OvDE-2 raw acoustic data

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The 10th International Workshop on Acoustic  
and Radio EeV Neutrino Detection Activities  
(ARENA 2024)

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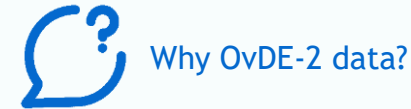
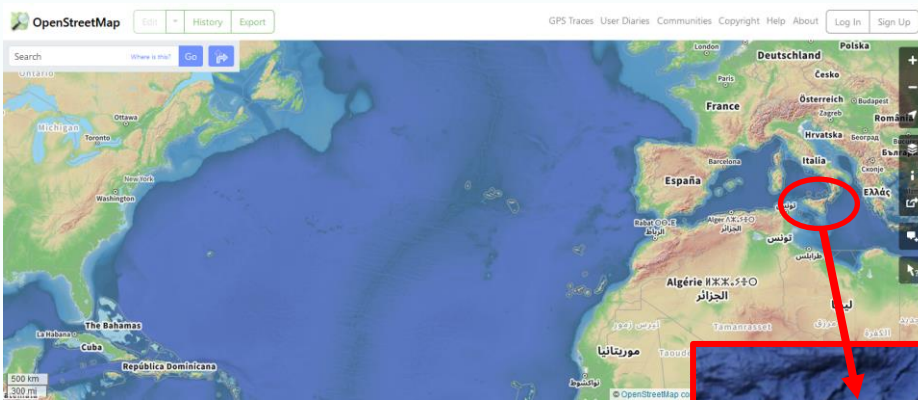
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## SUMMARY AND CONCLUSIONS

# INTRODUCTION

## OvDE-2

The aim of this presentation: Test an acoustic neutrino detector (presented at [ARENA22](#)) in another environment. The data is provided by a hydrophone in OvDE-2 (2100 m depth).



OvDE-2 is a marine station with four SMID TR-401 hydrophones + preamplifiers (synchronized by a GPS signal)  
 $f_s: 192 \text{ kHz}$  (24 bits)



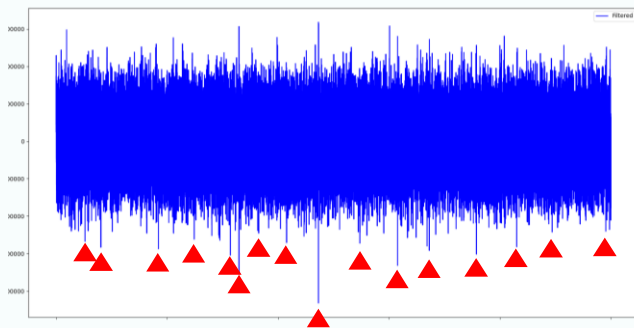
First studies for the experimental acoustic neutrino detection in the Ionian Sea has been done by the former NEMO Collaboration with the deployment of OvDE (Ocean Noise Detection Experiment).

The construction of OvDE-2 has represented a test-bench for the SMO (Submarine Multidisciplinary Observatory) project.

# INTRODUCTION

## OvDE-2

The aim of this presentation: Test an acoustic neutrino detector (presented at [ARENA22](#)) in another environment. The data is provided by a hydrophone in OvDE-2 (2100 m depth).

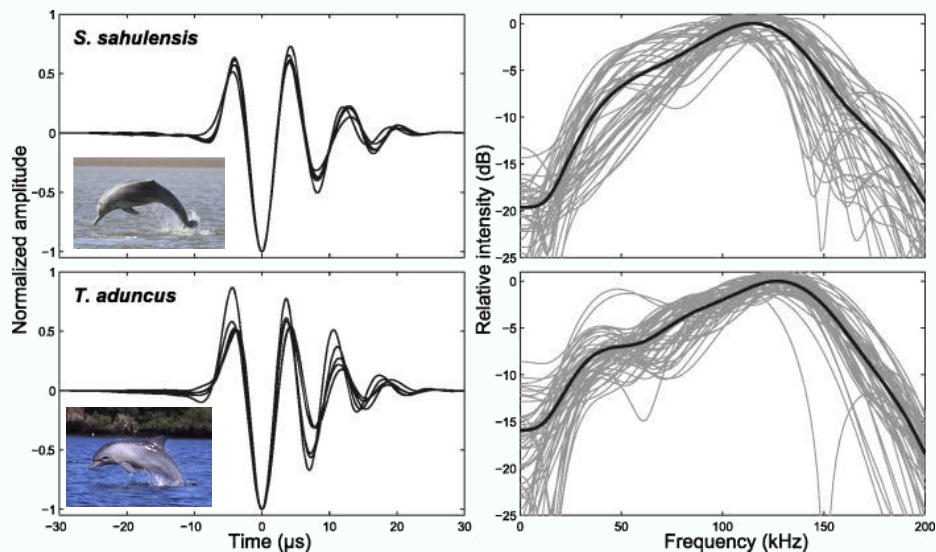


▲ The worst case: a lot of bioacoustics clicks!

We want to test the acoustic neutrino trigger in a critical case environment: 24H recorded by OvDE-2 with a lot of bioacoustics clicks for the marine mammals' echolocation.



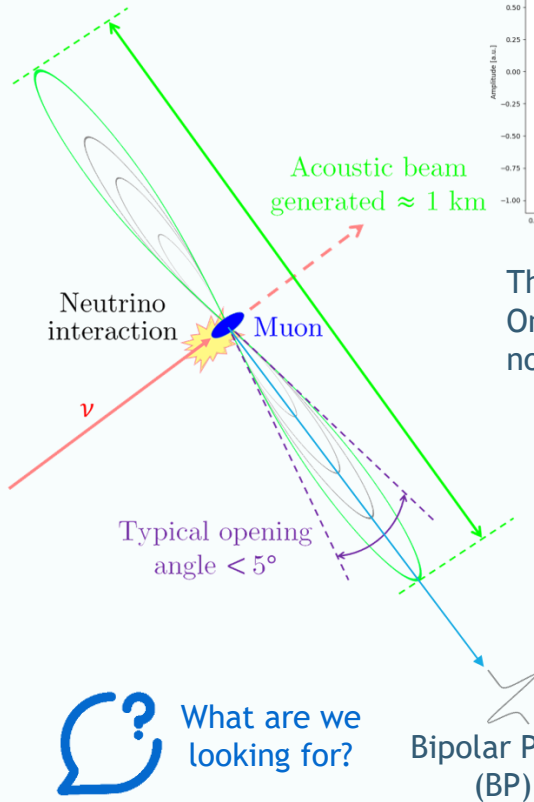
How sensitive is the detector to these clicks?



[M.. De Freitas, et.al Echolocation parameters of Australian humpback dolphins (*Sousa saahulensis*) and Indo-Pacific bottlenose dolphins (*Tursiops aduncus*) in the wilde. *The Journal of the Acoustical Society of America*. 2015, 137, 3033–3041. DOI: [10.1121/1.4921277](https://doi.org/10.1121/1.4921277)]

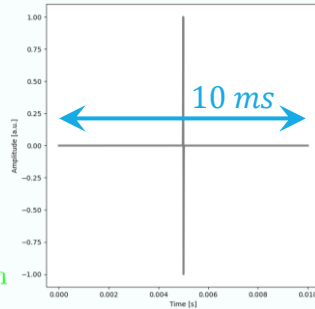
# INTRODUCTION

## NEUTRINO INDUCED BIPOLAR PULSE



What are we looking for?

BP generation at 1 MHz



UHE Neutrinos (peak):

- $10^{12} \text{ GeV} \approx 1.25 \text{ Pa}$
- $10^{11} \text{ GeV} \approx 123.9 \text{ mPa}$
- $10^{10} \text{ GeV} \approx 12.33 \text{ mPa}$
- ...



[S. Bevan, et al Study of the acoustic signature of UHE neutrino interactions in water and ice. *Nucl. Instrum. Methods Phys. Res., Sect. A* **2009**, 607, 398–411. DOI: [10.1016/j.nima.2009.05.009](https://doi.org/10.1016/j.nima.2009.05.009)]

OvDE-2  $f_s = 192 \text{ kHz}$

There are several equations to simulate a BP from the interaction of a neutrino in fluid. On this occasion, it has been chosen to simplify the generation of a BP (t): by deriving a normal distribution  $g(t)$

$$g(t) = e^{-\frac{1}{2}\left(\frac{t}{\sigma}\right)^2}$$

where the  $\sigma$  represents the standard deviation at a confidence interval of  $\sim 68\%$ . To control the width of the BP, the inter-peak value  $\Lambda$  should be equivalent to  $2\sigma$ :

$$BP(t) = \frac{dg}{dt} = -\frac{t}{\sigma^2} \cdot e^{\left[\frac{-t^2}{2\sigma^2}\right]}$$

$$BP(t) = \frac{-4t}{\Lambda^2} \cdot e^{-2\left(\frac{t}{\Lambda}\right)^2}$$



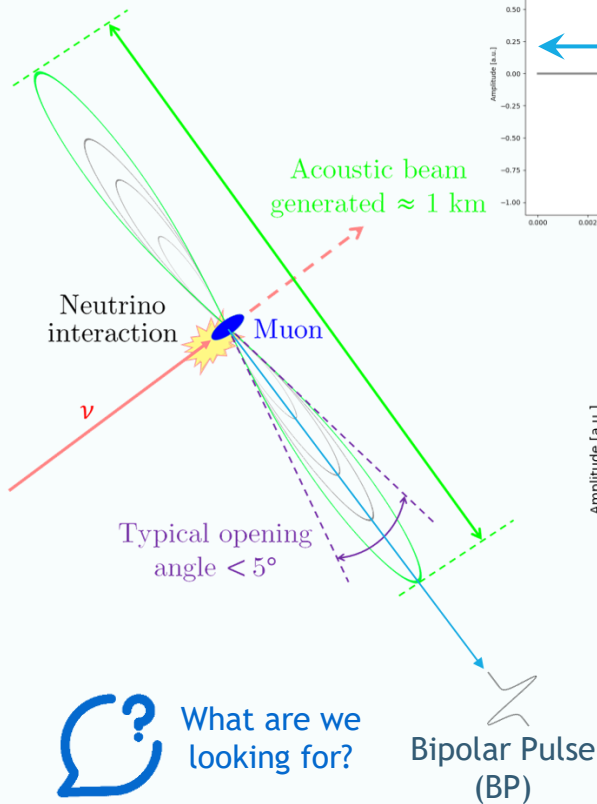
$$\Lambda = 2\sigma$$

$\Lambda$  represents the BP inter-peak time distance

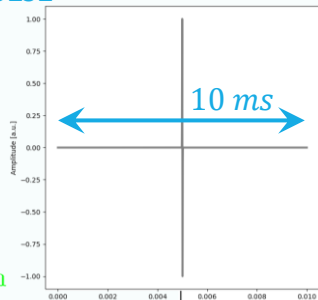
The acoustic neutrino signature is an **exotic signal** that contains a range of low frequencies ( $< 100 \text{ kHz}$ ) and it is propagated in a very narrow beam ( $< 5^\circ$ )

# INTRODUCTION

## NEUTRINO INDUCED BIPOLAR PULSE



BP generation at 1 MHz

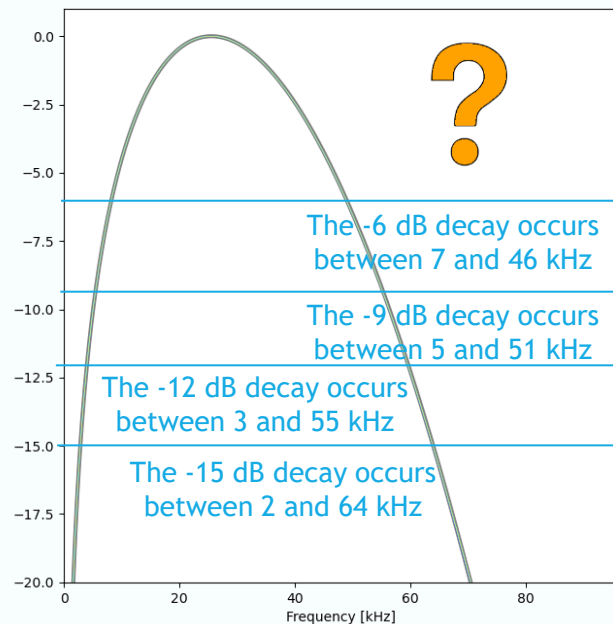
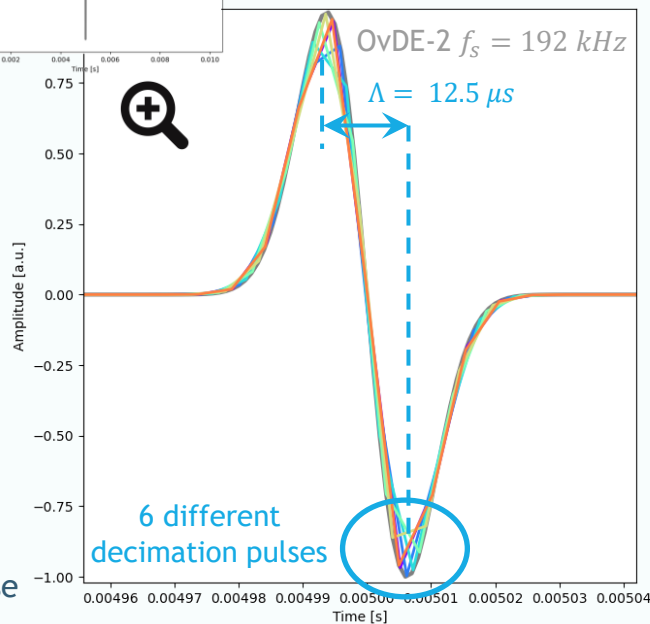


UHE Neutrinos (peak):

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Simulated BP at  $0^\circ$  (hydrophone  $\perp$  pancake) and 1 km distance

# ACOUSTIC NEUTRINO DETECTION

The background features a white space on the left and a complex geometric pattern of overlapping blue triangles on the right. The triangles vary in shades from light sky blue to deep navy blue, creating a dynamic, layered effect.

# ACOUSTIC NEUTRINO DETECTION

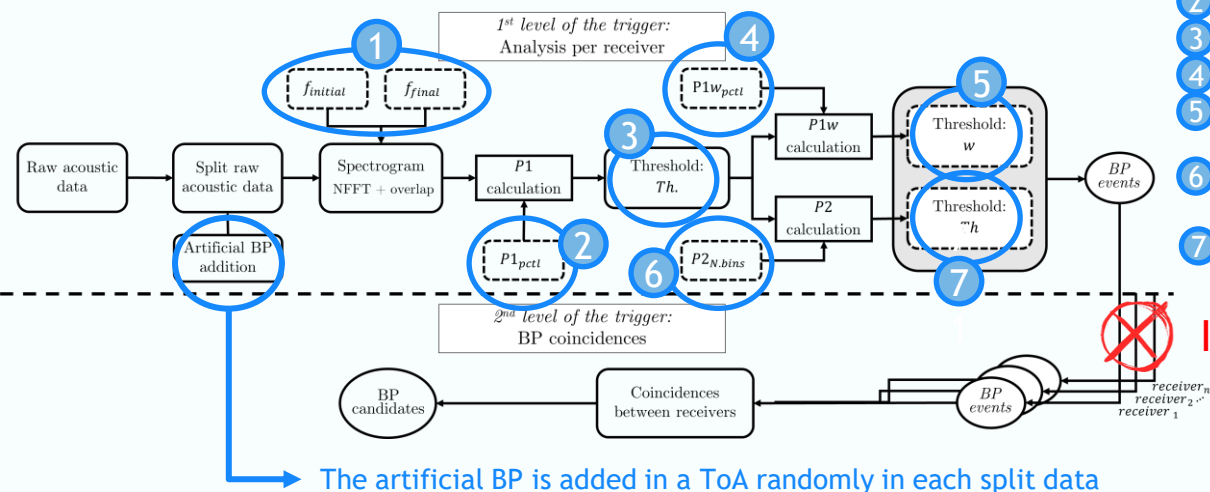
## WORKFLOW

The proposed acoustic neutrino detector at [ARENA22](#) studies 3 parameters (cut-off) provided by the spectrogram analysis:

- $P1$  (to control the BP intensity): Mean PSD value between the BP frequency range
- $P1w$  (to control the BP duration): Width to  $P1_{pctl}$ % of the  $P1$  peak
- $P2$  (to control the BP duration): Difference between  $P1$  value of 1 bin ( $\sim 50 \mu s$ ) and the  $P1$  value of the  $\pm 50$  samples surrounding it ( $\sim 5 ms$ )

### Parameters set-up:

- 1  $f_{initial} + f_{final}$ : Frequency range
- 2  $P1_{pctl}$ : Percentile for the  $P1$  calculation
- 3  $P1_{Th}$ : Threshold of  $P1$  to detect BP events
- 4  $P1w_{pctl}$ : Percentile for the  $P1w$  calculation
- 5  $P1w_{Th}$ : Threshold of  $P1w$  to detect BP events
- 6  $P2_{N.bins}$ : Bins surrounding the candidate for the  $P2$  calculation
- 7  $P2_{Th}$ : Threshold of  $P2$  to detect BP events



**In this study, we are using a single hydrophone**



# ACOUSTIC NEUTRINO DETECTION

## ARTIFICIAL BP IN OVDE-2 RAW DATA

The inserted BP is at random time of the raw acoustic data



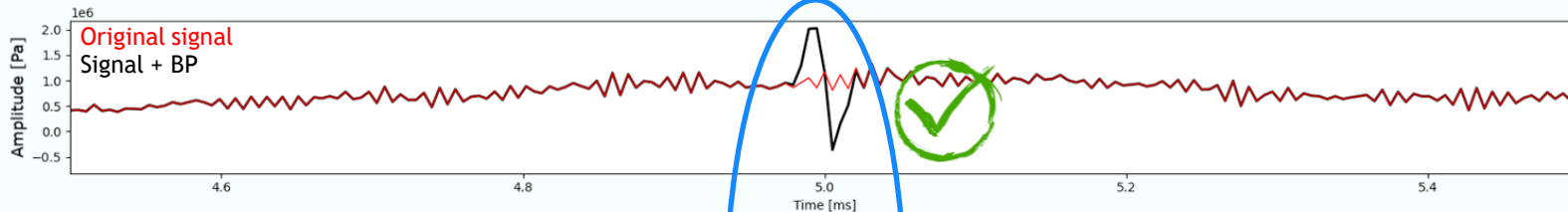
Which neutrinos are “visible” in OvDE-2?

UHE Neutrinos (peak):

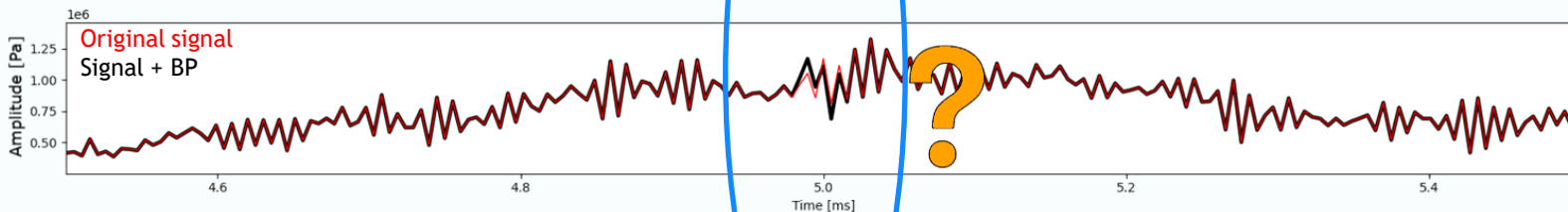
- $10^{12} \text{ GeV} \approx 1.25 \text{ Pa}$
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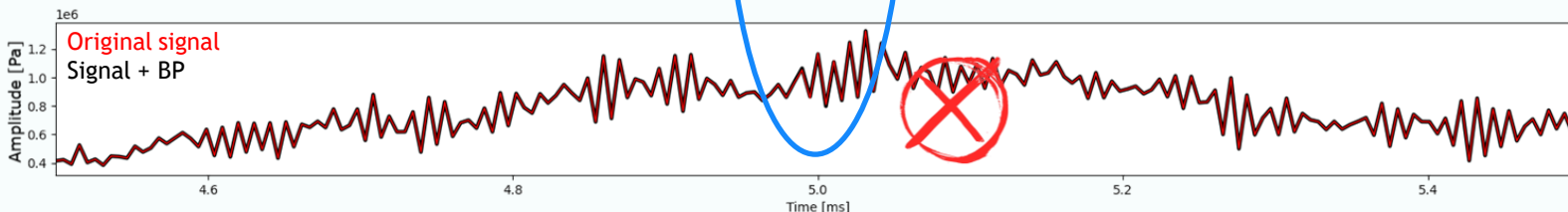
$10^{12} \text{ GeV}$



$10^{11} \text{ GeV}$



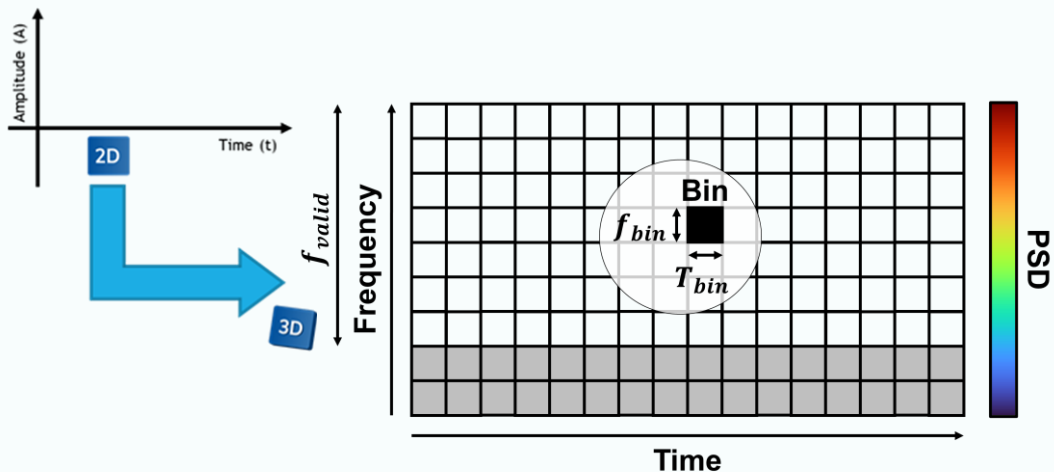
$10^{10} \text{ GeV}$



The spectrogram analysis is proposed because the detection depends more on energy than BP shape

# ACOUSTIC NEUTRINO DETECTION

## SPECTROGRAM CONFIGURATION



$P1$ : Mean PSD value between the BP frequency range

$P1w$ : Width to  $P1_{pctl}\%$  of the  $P1$  peak

$P2$ : Difference between  $P1$  value of 1 bin and the  $P1$  value of the  $\pm 50$  bins surrounding it

$f_s$ : 192 kHz

Compromise between  $f_{valid}$ ,  $T_{bin}$ ,  $f_{res}$  and  $f_s$



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D. Bonanno, L. Di Mauro, A. Idrissi, G. Riccobene, S. Sanfilippo, S. Viola

$$T_{bin} = \frac{NFFT}{f_s} (1 - overlap)$$

$$f_{bin} \geq \frac{f_s}{NFFT}$$

$$f_{valid} \geq \frac{2f_s}{NFFT}$$

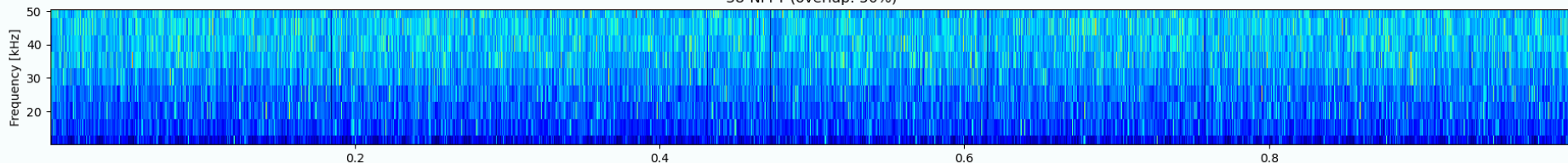
NFFT	$T_{bin}$ [ $\mu$ s]	$f_{bin}$ [kHz]	$f_{valid}$ [kHz]
	Objective $\lesssim 100$	Objective $\lesssim 5$	Objective $\lesssim 10$
32	83.33	6.00	12.00
38	98.96	5.05	10.10
64	166.67	3.00	6.00

# ACOUSTIC NEUTRINO DETECTION

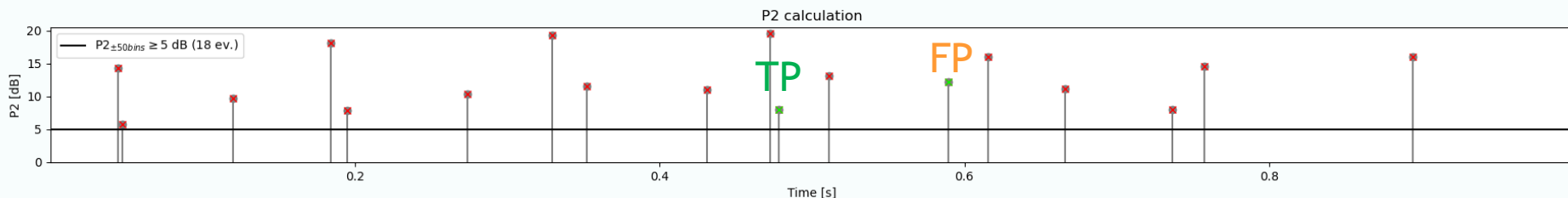
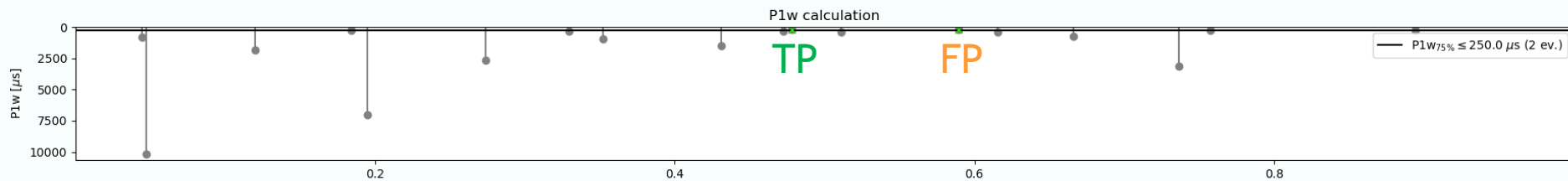
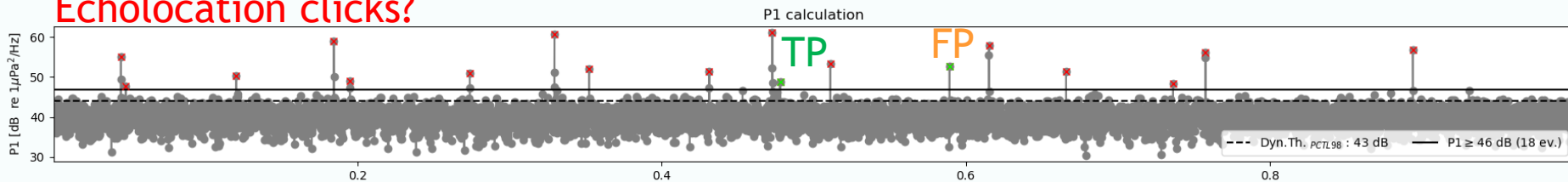
TP: True Positive  
FP: False Positive

## THE DETECTOR OPERATION

Frequencies for the detection: 10.1-50.5 kHz  
38-NFFT (overlap: 50%)



## Echolocation clicks?

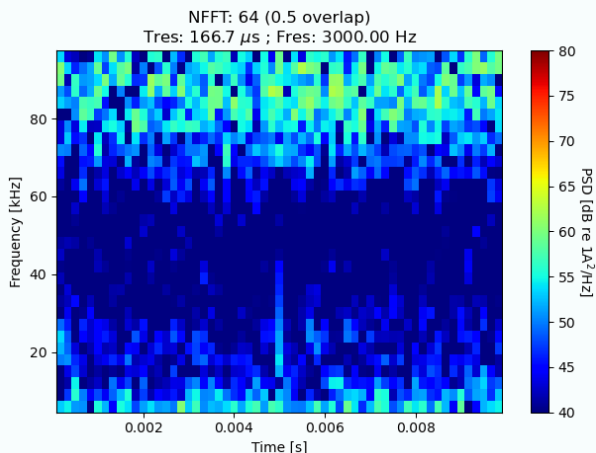
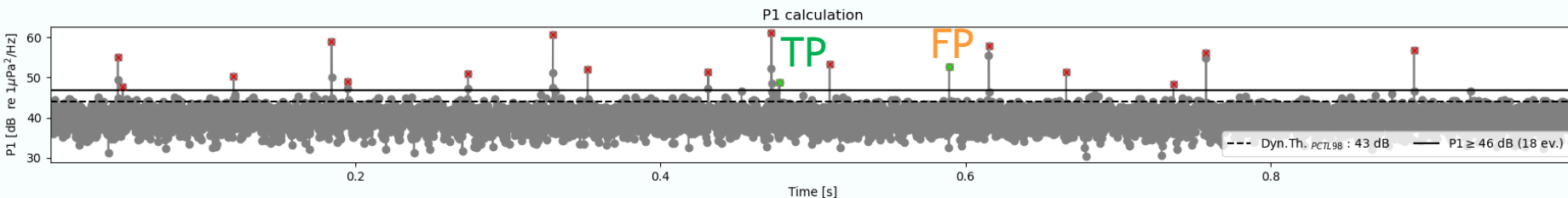
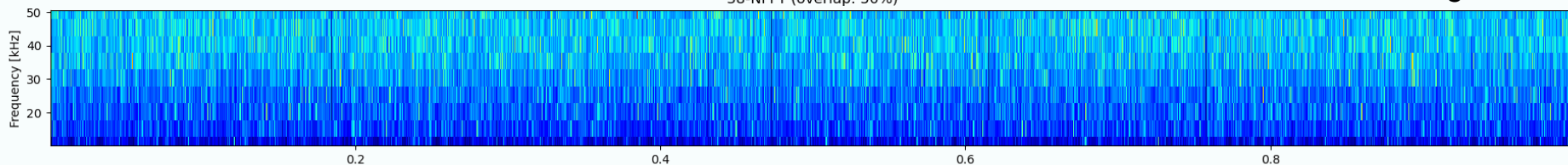


# ACOUSTIC NEUTRINO DETECTION

## THE DETECTOR OPERATION

TP: True Positive  
 FP: False Positive  
 FN: False Negative

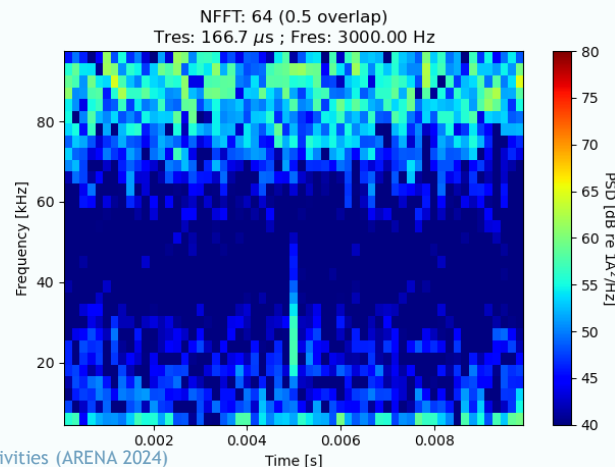
Frequencies for the detection: 10.1-50.5 kHz  
 38-NFFT (overlap: 50%)



TP  
 Powerful and short clicks  
 pass the algorithm as a BP

$$Precision = \frac{TP}{TP + FP}$$

$$Recall = \frac{TP}{TP + FN}$$

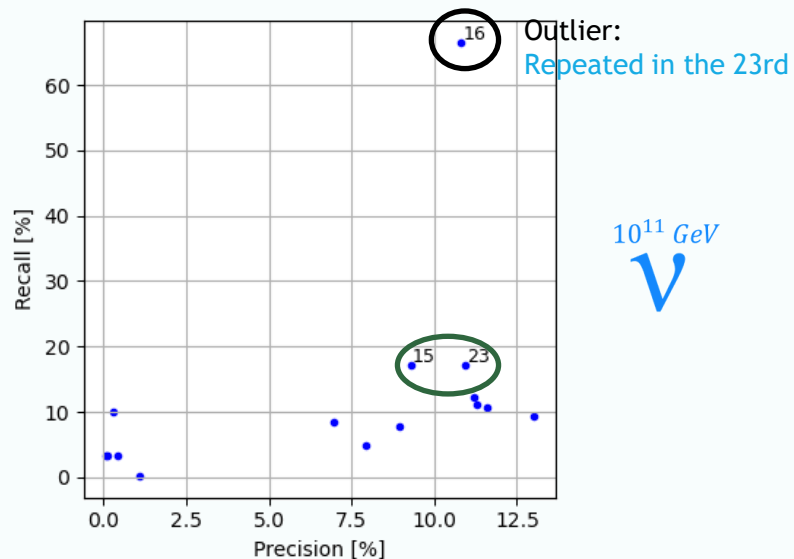


# ACOUSTIC NEUTRINO DETECTION

## THE DETECTOR CONFIGURATION

Different tests have been carried out to identify the optimal detector configuration. A random hour of the total 24H sample has been used.

We proceeded to test different frequency ranges, different values for the parameters calculation, and different thresholds.

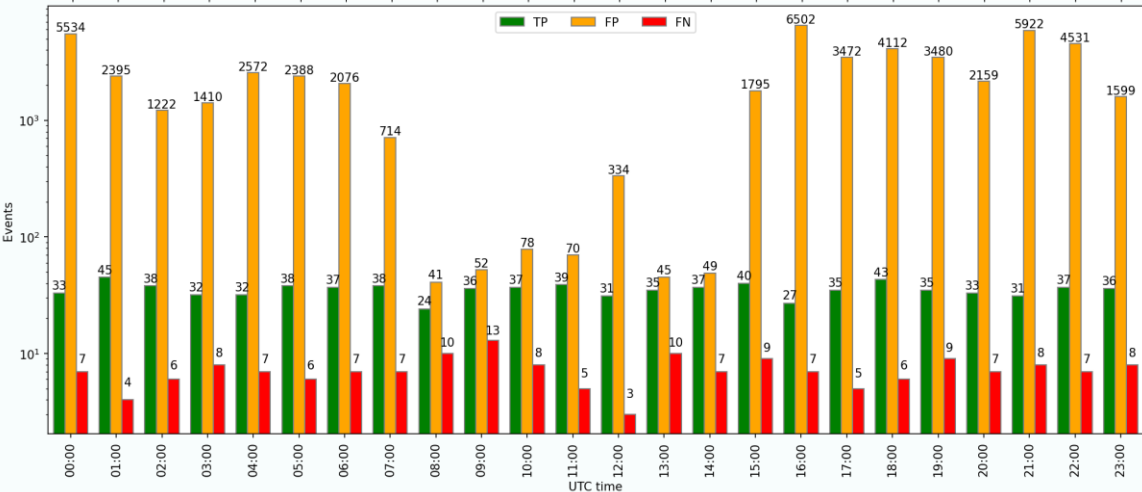
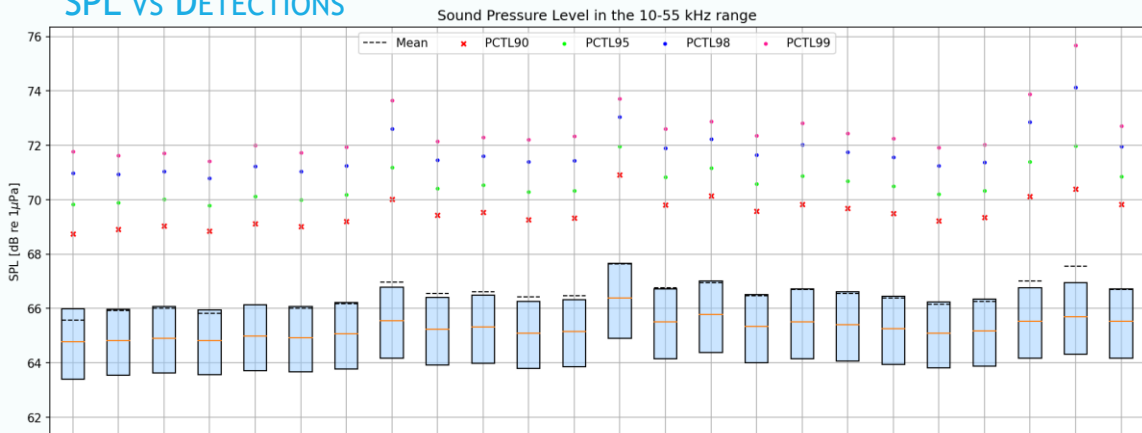


Index	Frequency range		P1 [dB re $1\mu\text{Pa}^2/\text{Hz}$ ]		P1w [s]	P2 [dB]			Precision [%]	Recall [%]
	Initial [kHz]	Final [kHz]	Percentile [%]	Threshold [dB]	Percentile [%]	Threshold [ $\mu\text{s}$ ]	Number of bins [ $\pm$ samples]	Threshold [dB]		
15	10	50	98	3	80	250	50	5	9.30	17.22
16	10	55	98	3	75	275	50	5	10.82	66.52
23	10	55	98	3	75	275	50	5	10.92	17.22
...										

# RESULTS

# RESULTS

## SPL VS DETECTIONS



Is the background noise in the raw data affecting the BP detector?

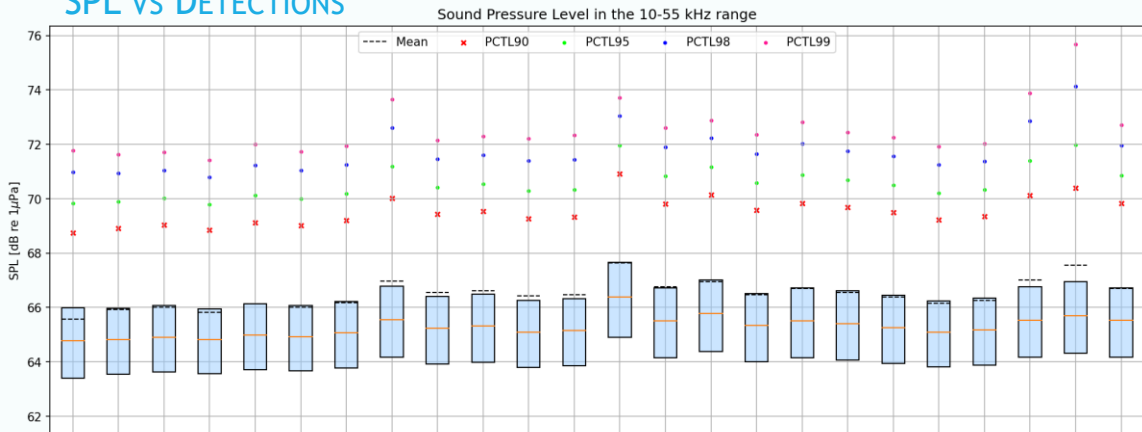
$10^{12}$  GeV  
V

TP detections: Constant  
FP detections: +SPL  $\rightarrow$  -FPs

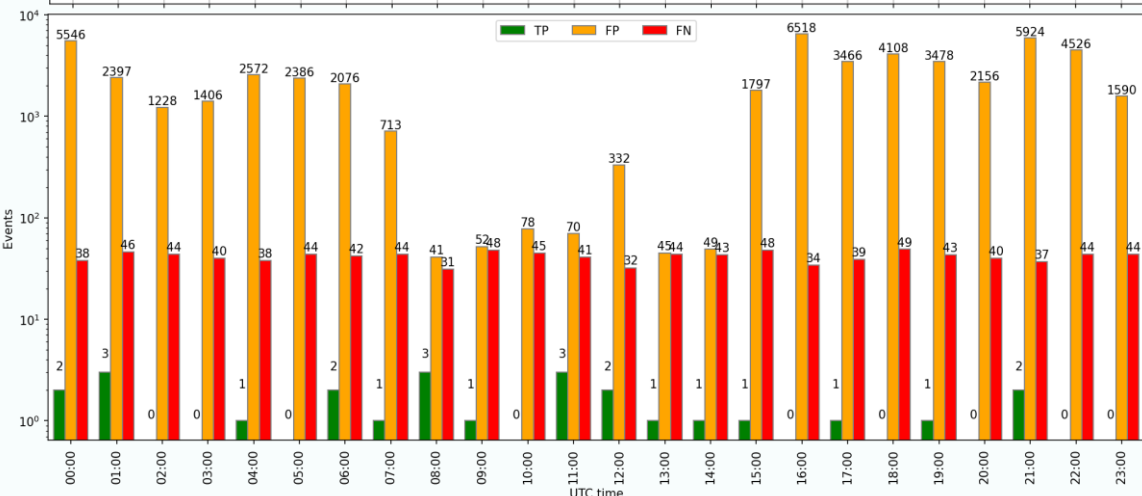
Precision: 1.59 %  
Recall: 82.99 %

# RESULTS

## SPL VS DETECTIONS



Is the background noise in the raw data affecting the BP detector?



$10^{11}$  GeV  
V

TP detections: Constant?

FP detections: +SPL  $\rightarrow$  -FPs

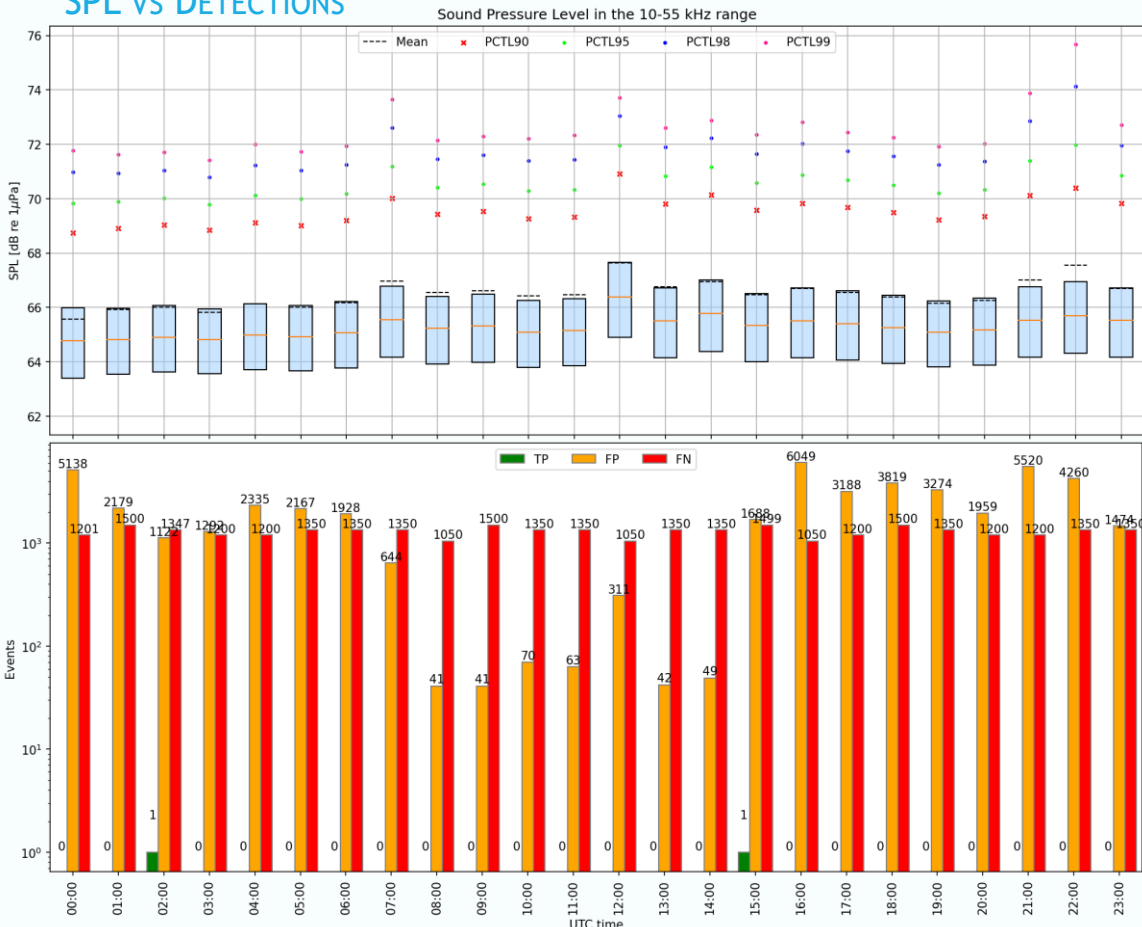
Precision: 0.5 %

Recall: 2.44 %



# RESULTS

## SPL vs DETECTIONS



Is the background noise in the raw data affecting the BP detector?

$10^{10}$  GeV

TP detections: 2 🤖

FP detections: +SPL  $\rightarrow$  -FPs

Precision: 0.041 %

Recall: 0.064 %

# RESULTS

## DETECTIONS

Energy [GeV]	Artificial BPs	BP detections	True Positive (TP)	False Positive (FP)	False Negative (FN)	Precision [%]	Recall [%]
$10^{12}$	1023 (1/min)	53401 (0.86/s)	849	52552	174	1.59	82.99
$10^{11}$	1023 (1/min)	52579 (0.84/s)	25	52554	998	0.05	2.44
$10^{10}$	1023 (1/min)	0 (0.00/s)	0	0	1023	0.00	0.00
$10^{10}$	31199 (30/min)	48655 (0.78/s)	2	48653	31197	0.0041	0.0064

Objective < 1/s



# RESULTS

## DETECTIONS

Precision: 1.59%

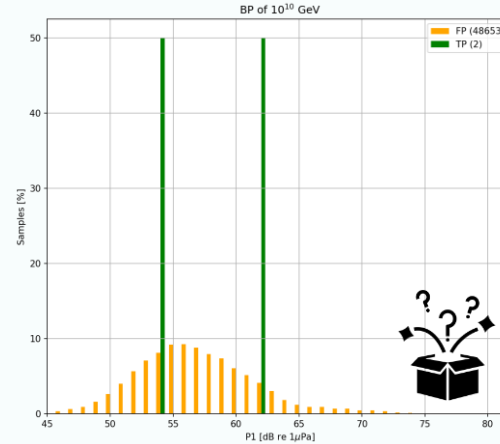
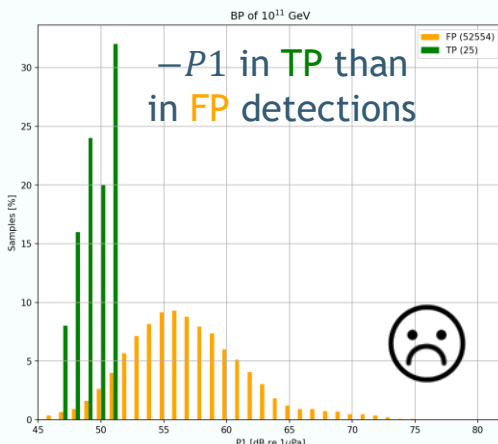
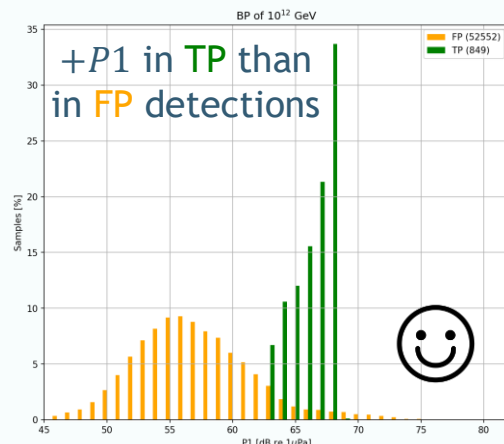
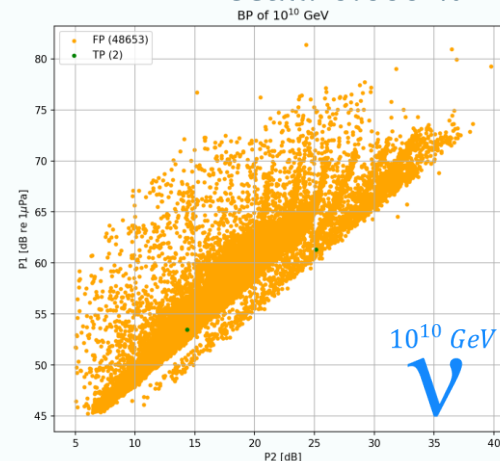
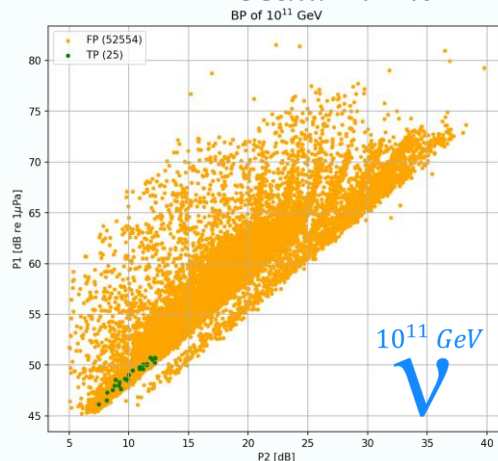
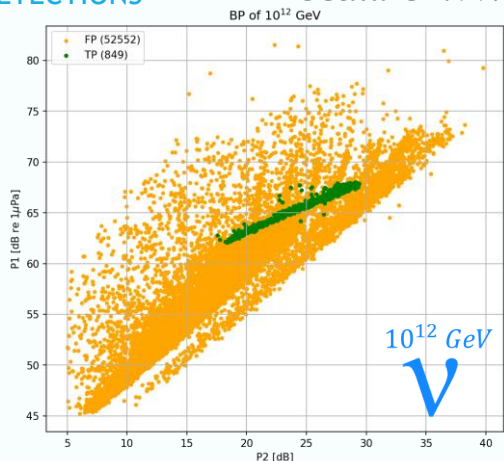
Recall: 82.99%

Precision: 0.05%

Recall: 2.44%

Precision: 0.0041%

Recall: 0.0064%



# SUMMARY AND CONCLUSIONS

- The algorithm proposed in [ARENA22](#) has been tested in another environment (critical case full of bioacoustics echolocation clicks).
- Neutrinos of energy  $10^{12}$  GeV we are above background noise (FPs).
- OvDE-2 data requires a different kind of detector (scalogram analysis?).
- Testing the second step of the algorithm (coincident events at different receivers) is expected to reduce the number of FPs.
- Research should be undertaken to develop detectors for lower energy neutrinos.
- This kind of detector can also be useful for other signals, for example for bioacoustics echolocation clicks.

# Thanks for your attention

The 10th International Workshop on Acoustic  
and Radio EeV Neutrino Detection Activitie  
(ARENA 2024)



# Backup slides

The 10th International Workshop on Acoustic  
and Radio EeV Neutrino Detection Activitie  
(ARENA 2024)



# ACOUSTIC NEUTRINO DETECTION

## THE DETECTOR OUTPUT

File name	ToA [unixtime]	Is an artificial BP?	P1 [dB re $1\mu\text{Pa}^2/\text{Hz}$ ]	P1w [s]	P2 [dB]	N.P1	N.P1w	N.P2	SPLs [dB re $1\mu\text{Pa}$ ]
...	Detection time	True/False	Detection values			Number of events that have passed the cut-off threshold			SPL in the analyzed split

# ACOUSTIC NEUTRINO DETECTION

\* ↑temporal resolution : ↓frequency resolution

## THE SPECTROGRAM AS A SIGNAL DETECTOR

### P1 calculation:

Trigger alert of candidate BP using a spectrogram in KM3NeT:

