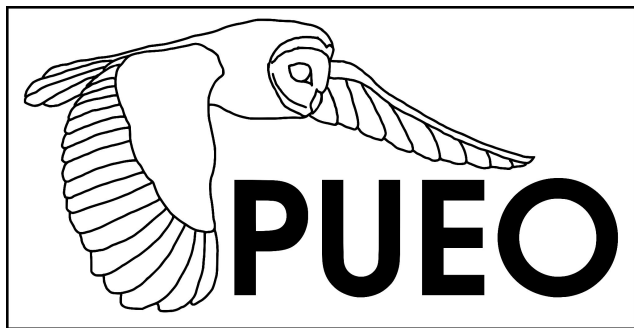


Prospects of Neutrino Flavor Measurements with PUEO

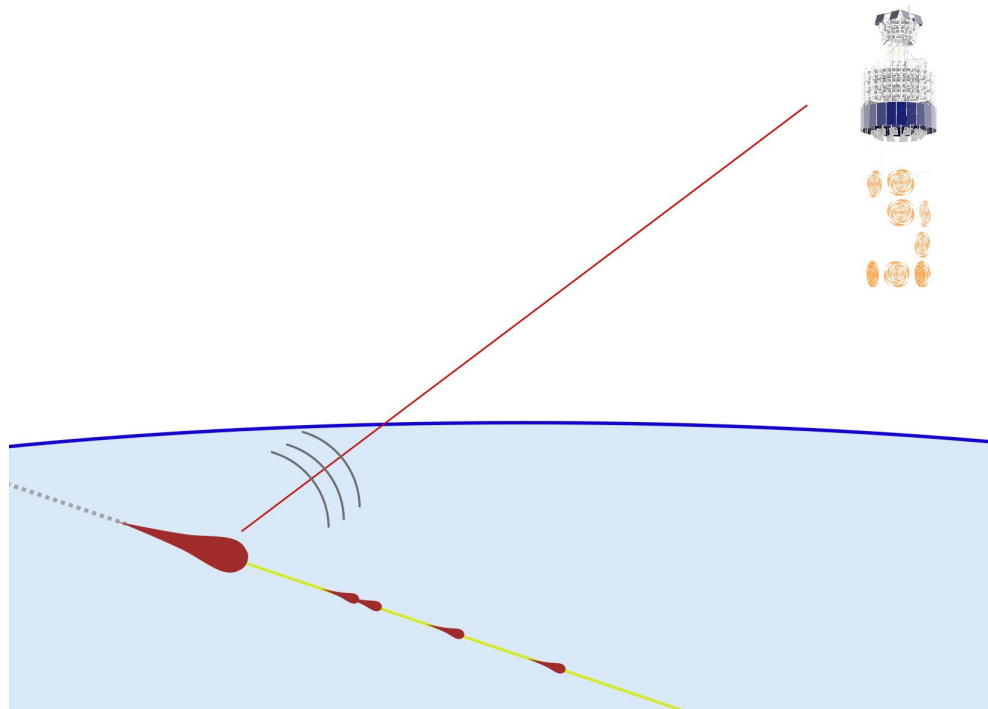
Christoph Welling, Austin Cummings, Rachel Scrandis
for the PUEO collaboration



Sensitivity to Flavor

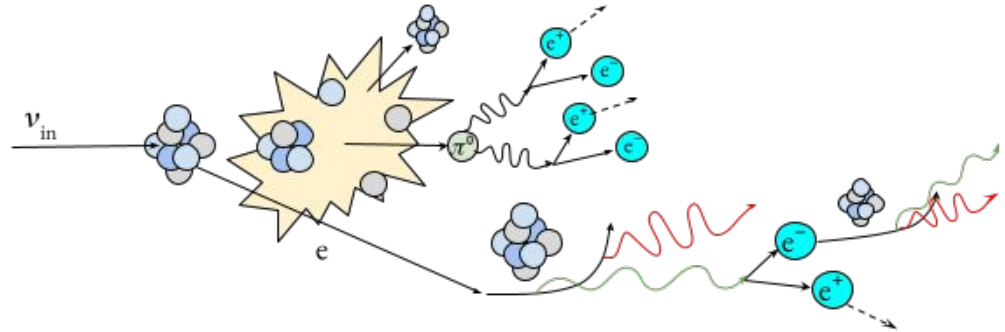
- Tau channel from LF instrument
- Secondary leptons from in-ice interactions
- ν_e :
 - Generates EM shower
- ν_μ and ν_τ :
 - μ or τ propagates
 - Secondary showers emit radio signals

PUEO should be able to see these!



Electron Neutrinos

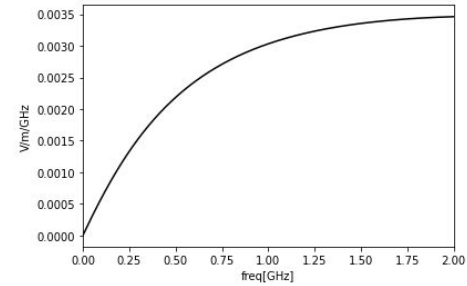
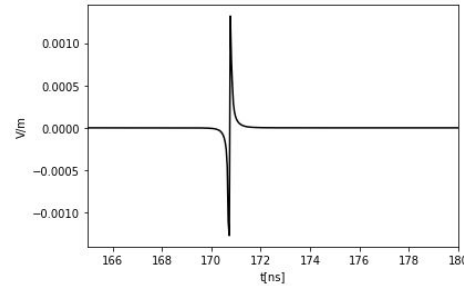
- EM shower on top of hadronic shower
- Offset between EM and hadronic shower maxima
- LPM effect: EM shower becomes longer and irregular
- Interference between radio signals from both showers



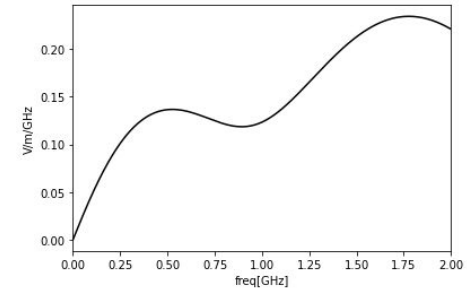
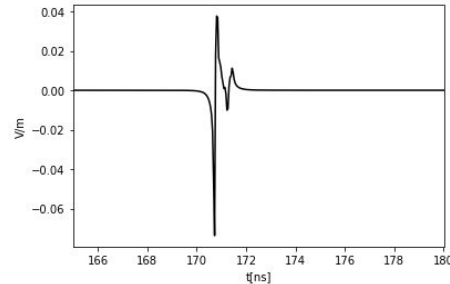
Electron Neutrinos

- EM shower on top of hadronic shower
- Offset between EM and hadronic shower maxima
- LPM effect: EM shower becomes longer and irregular
- Interference between radio signals from both showers

Hadronic shower

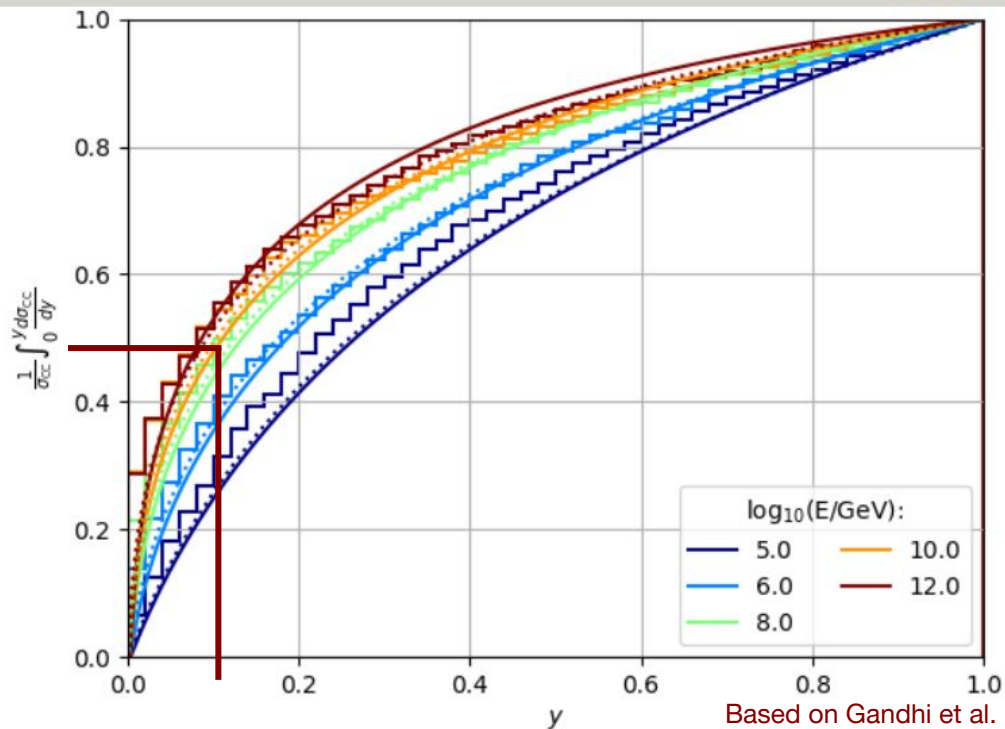


Hadronic + EM shower

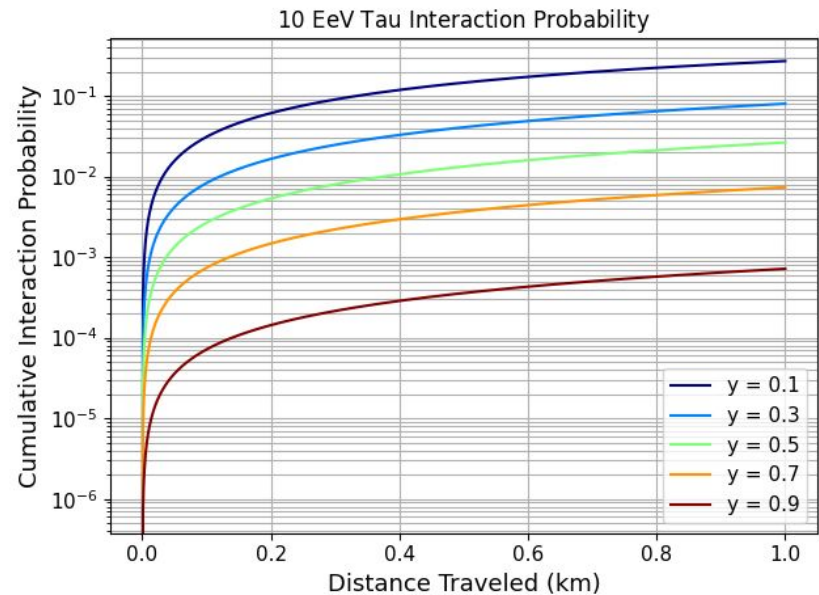
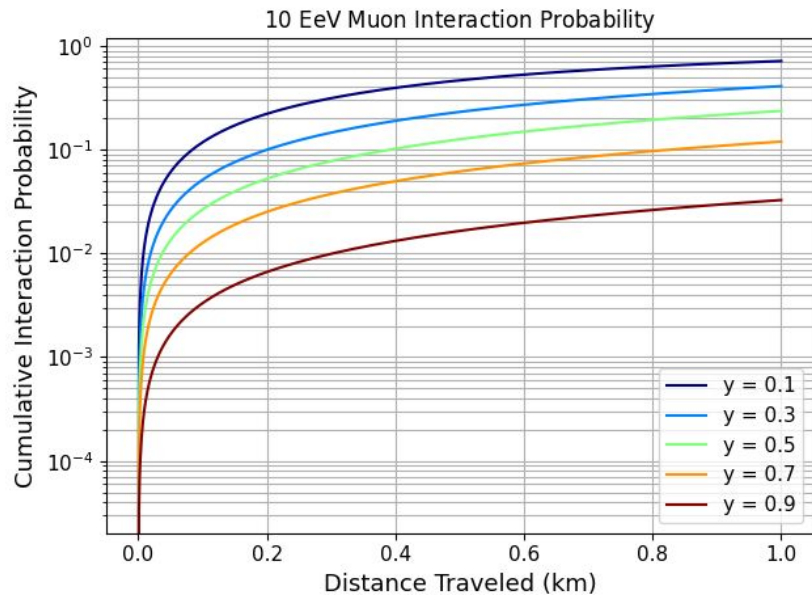


Muon and Tau Neutrinos

- Inelasticity distribution is heavily tilted towards low y
- Almost the time 90% or more of the energy goes into the mu/tau

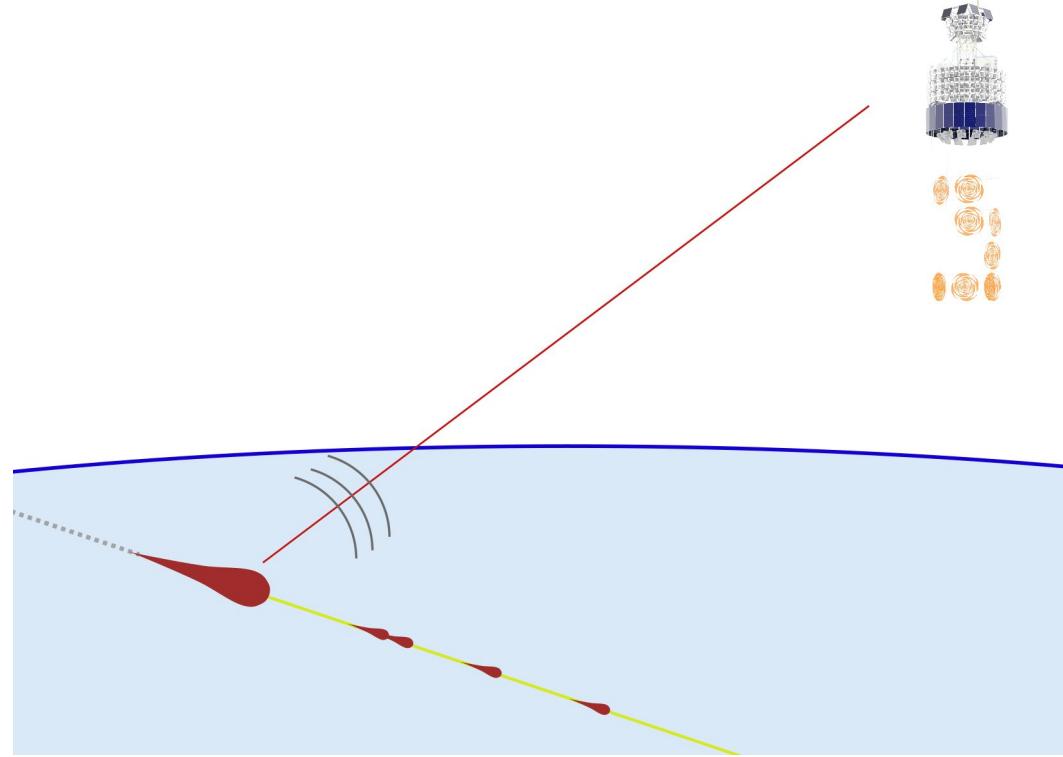


Muon and Tau Neutrinos



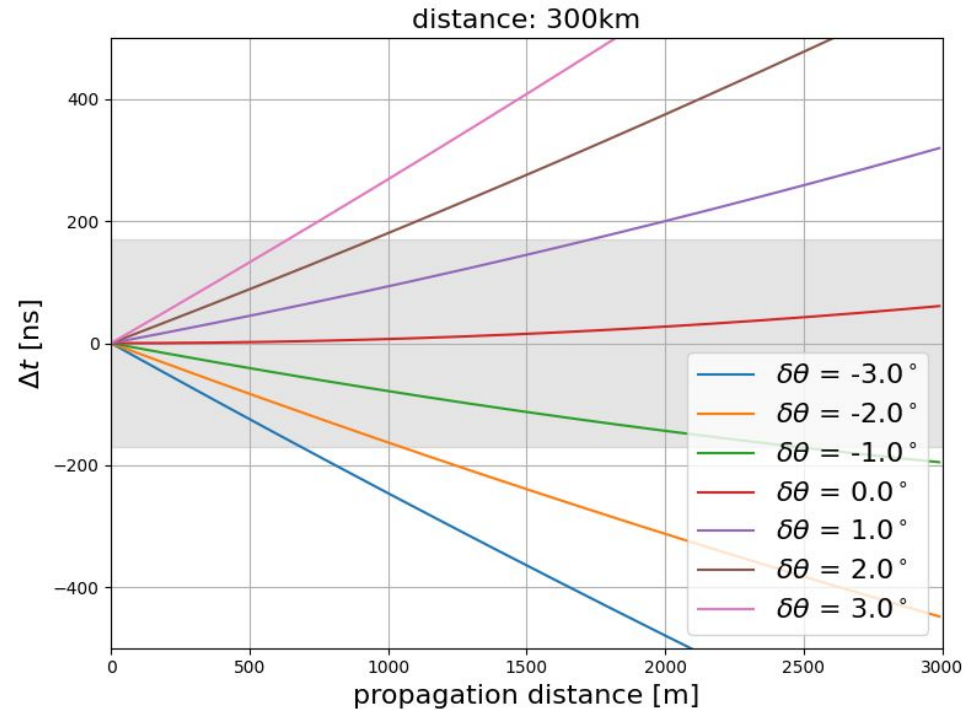
Radio Signals from Secondary Showers

- Events seen by PUEO will be $O(100\text{km})$ away
- Secondary showers are guaranteed to be seen at the Cherenkov angle!



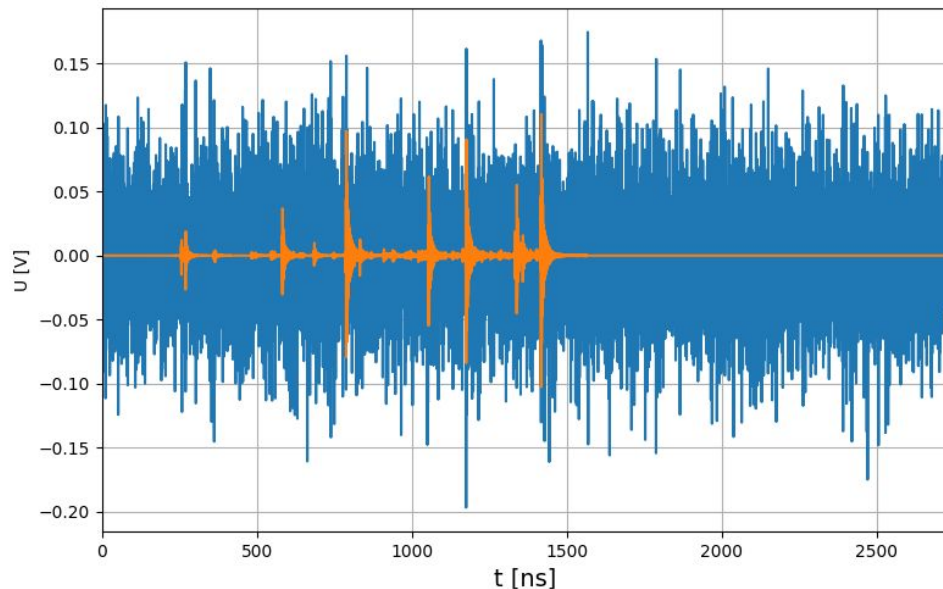
Time Compression Effects

- At the Cherenkov angle, radio emissions are compressed in time
- This applies to the time between showers too
- Secondary radio signals are likely to appear in recorded window
- This lets us look for sub-threshold events



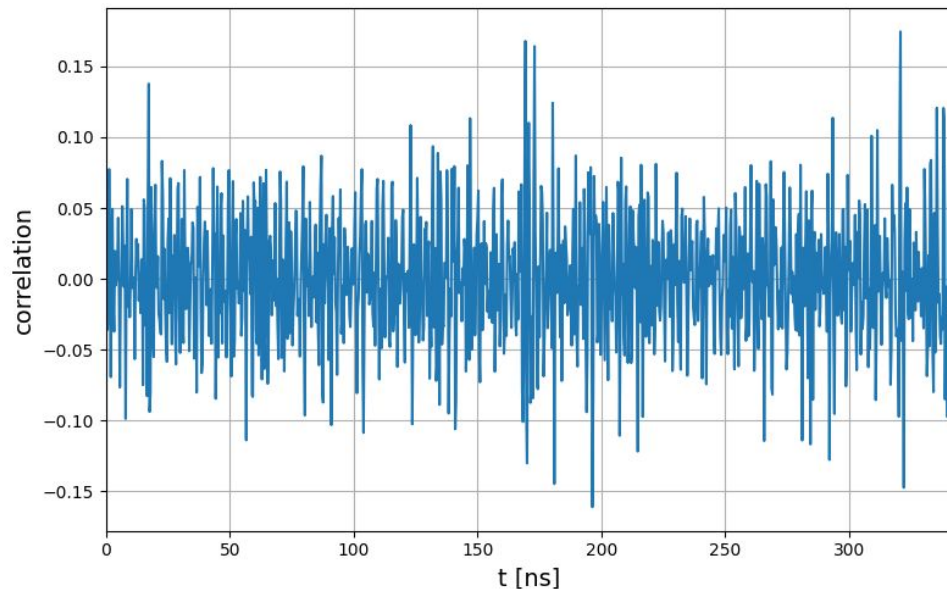
Secondary Shower Simulation

- Lepton propagation with NuLeptonSim [arxiv:2311.03646]
- Lepton is propagated until it enters atmosphere or bedrock
- Signal generation & propagation with NiceMC [PoS (ICRC2023)1154]
- Detector simulation with PueoSim



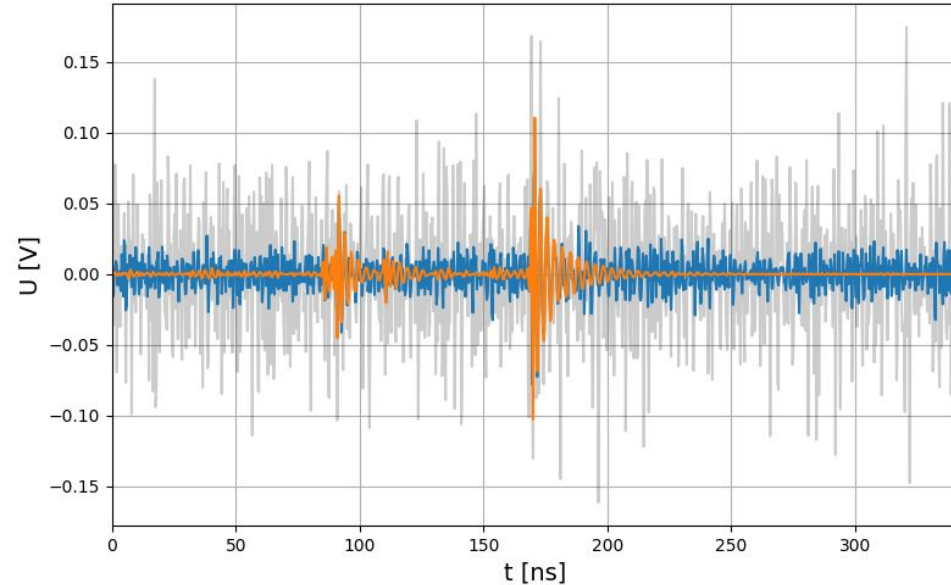
Secondary Shower Search

- Assumptions:
 - We triggered on the largest radio pulse in the waveform
 - We know this event is a neutrino
 - We know where the radio signal from the neutrino is
- Goal: Show that there are more showers in the recorded waveform



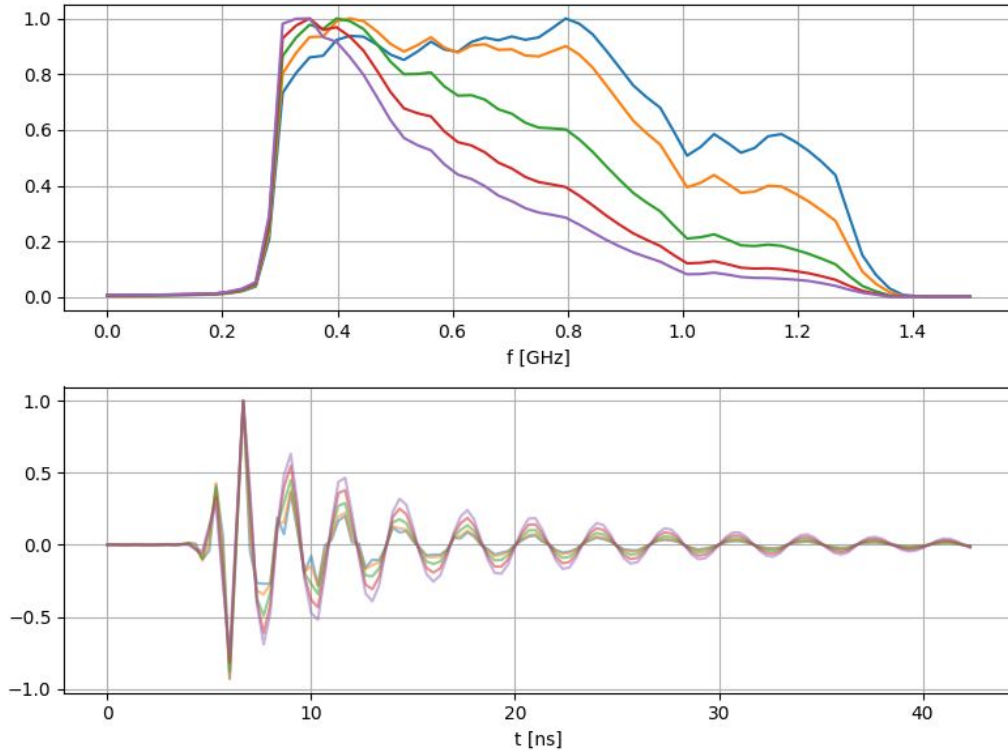
Beamforming

- Combine signals from 20 channels
- Time offsets determined from main RF pulse
- Choose between Vpol and Hpol channels



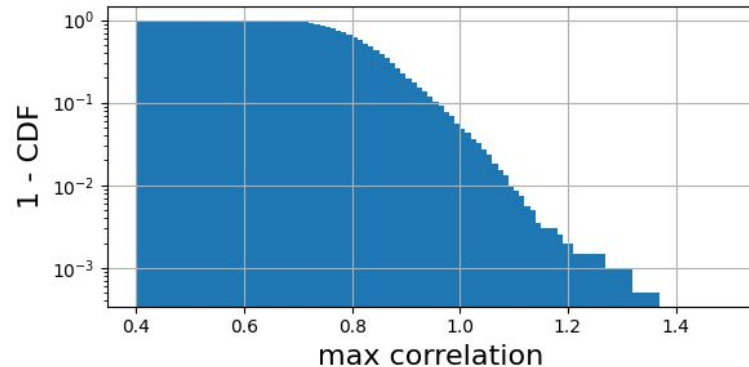
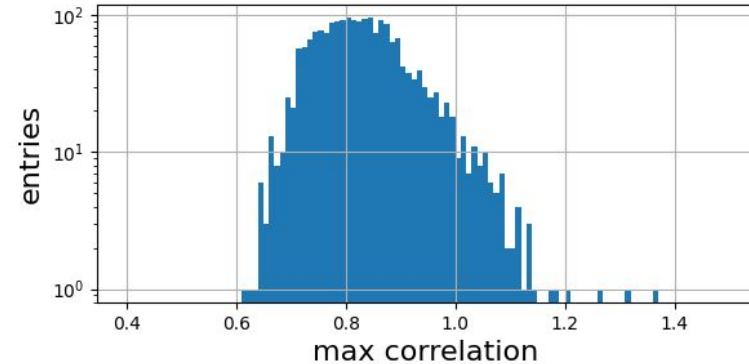
Templates

- Templates from Askaryan pulse folded with antenna & amplifier response
- Multiple templates for different viewing angles
- Choose template that best fits main RF pulse
- Background estimate from correlations with noise



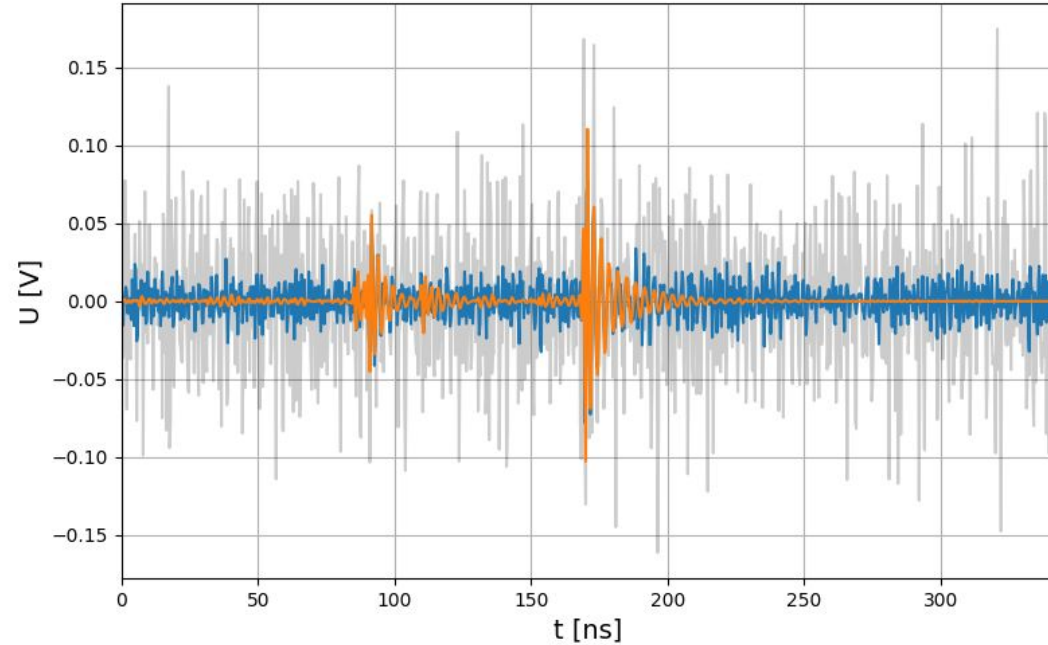
Templates

- Templates from Askaryan pulse folded with antenna & amplifier response
- Multiple templates for different viewing angles
- Choose template that best fits main RF pulse
- Background estimate from correlations with noise



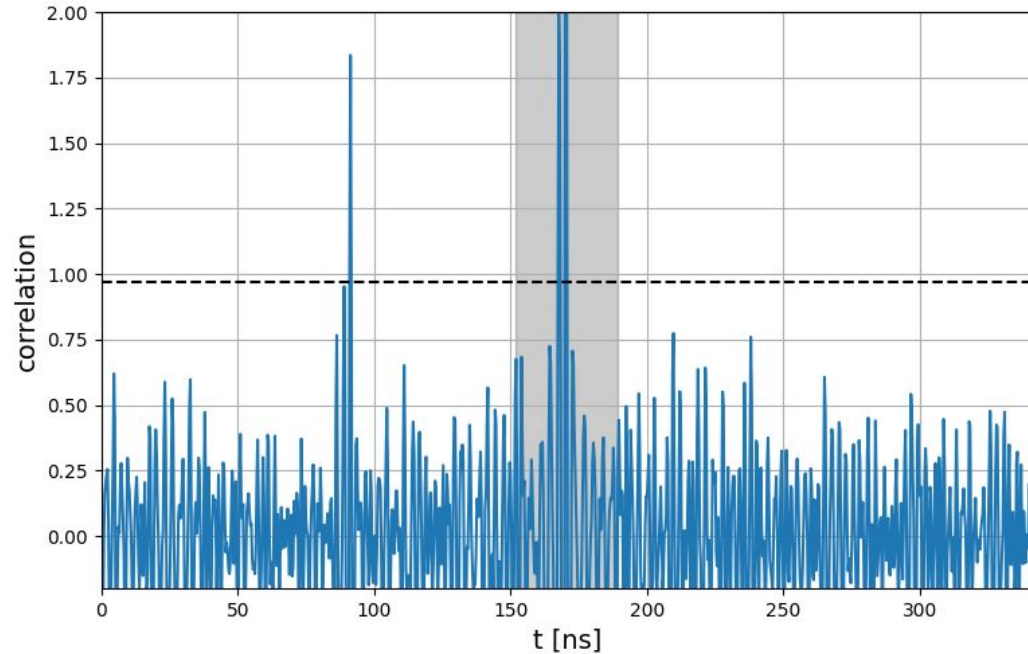
Template Correlation

- Correlate template with recorded waveform
- Mask region around main RF pulse
- Check of correlation is above x percentile



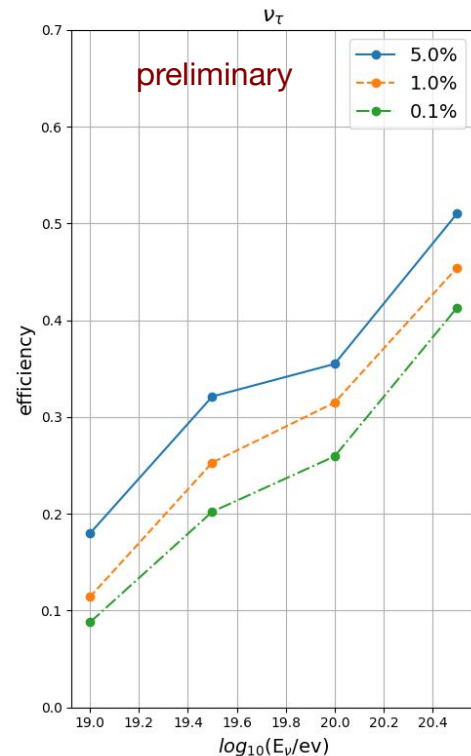
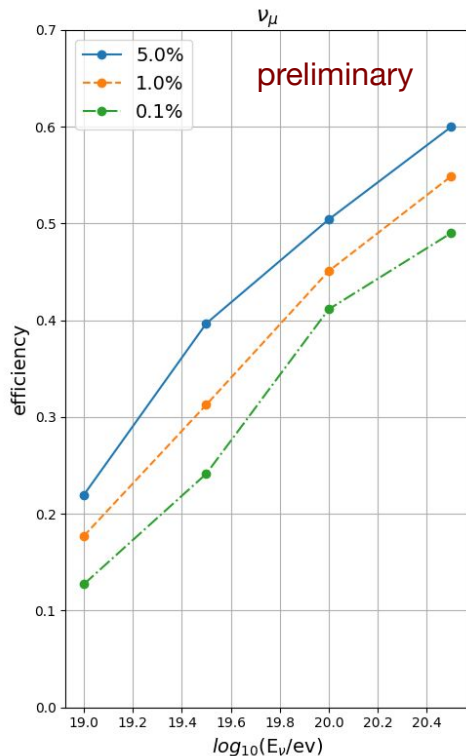
Template Correlation

- Correlate template with recorded waveform
- Mask region around main RF pulse
- Check of correlation is above x percentile



Efficiencies

- We can see a large fraction of secondary showers!
- Analysis still has a lot of room for improvement
- Double triggers could add to this further



Conclusions

- PUEO will have 3 channels sensitive to flavor:
 - Earth-skimming ν_τ through low frequency instrument
 - ν_e through signal spectrum
 - $\nu_\mu + \nu_\tau$ through secondary showers
- Potential to measure all 3 neutrino flavors
- Sensitivity to be determined, but first results look promising!

