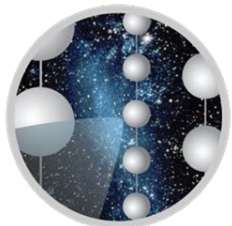


# Improving air-shower observations with the Surface Array Enhancement - Results from a Prototype Station

Megha Venugopal for the IceCube Collaboration

ARENA 2024, Chicago



ICECUBE

[megha.venugopal@kit.edu](mailto:megha.venugopal@kit.edu)

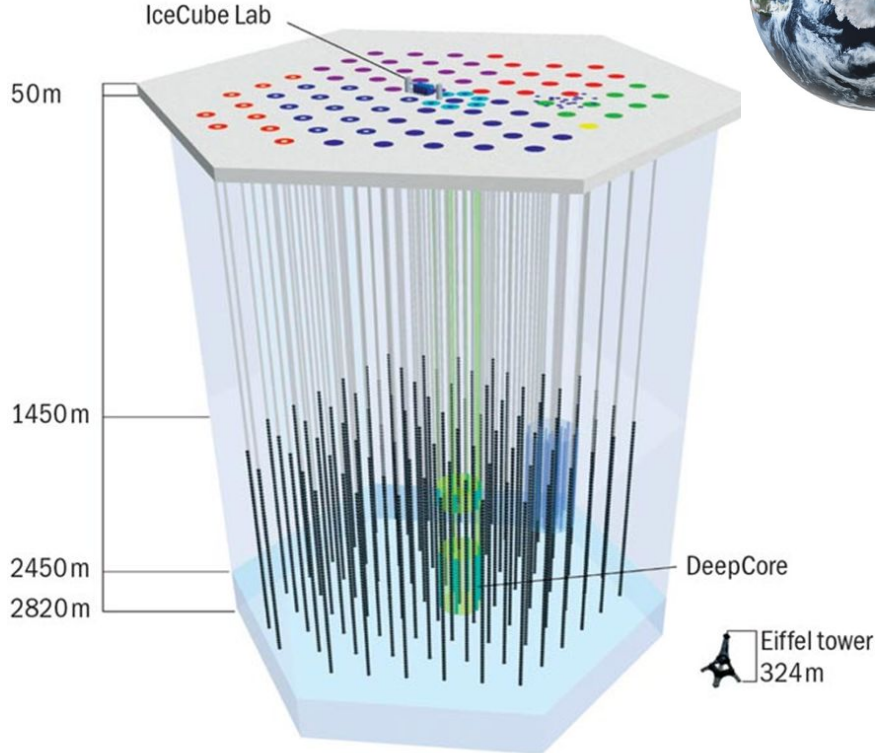


CHICAGO 2024

# IceCube and IceTop

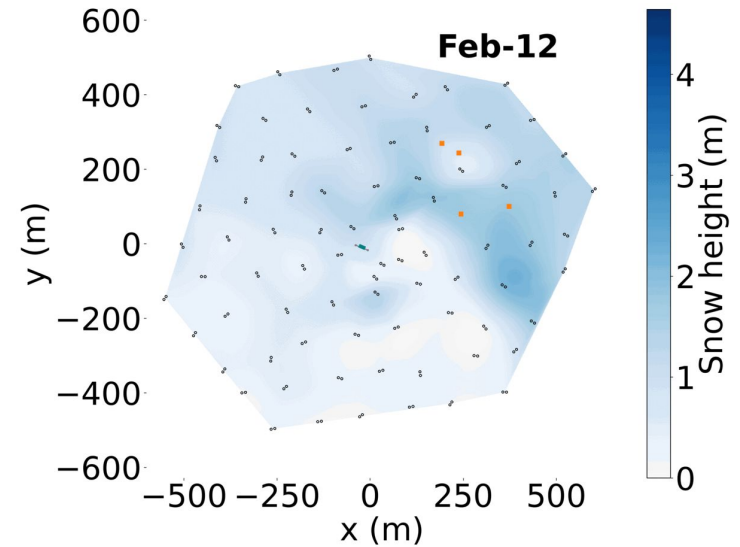
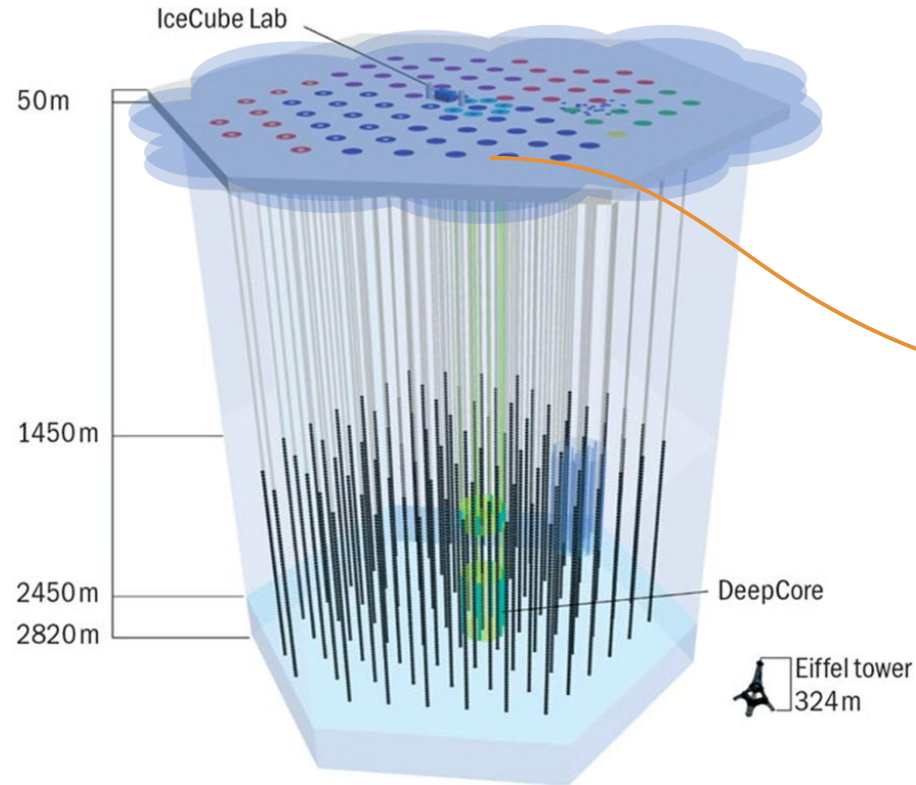


- IceCube is a  $\text{km}^3$  detector in ice detecting neutrinos and atmospheric muons
- IceTop is a  $\text{km}^2$  surface array for cosmic ray research and also serves as a veto for IceCube neutrino research.



Deployment of an IceTop tank

# IceTop

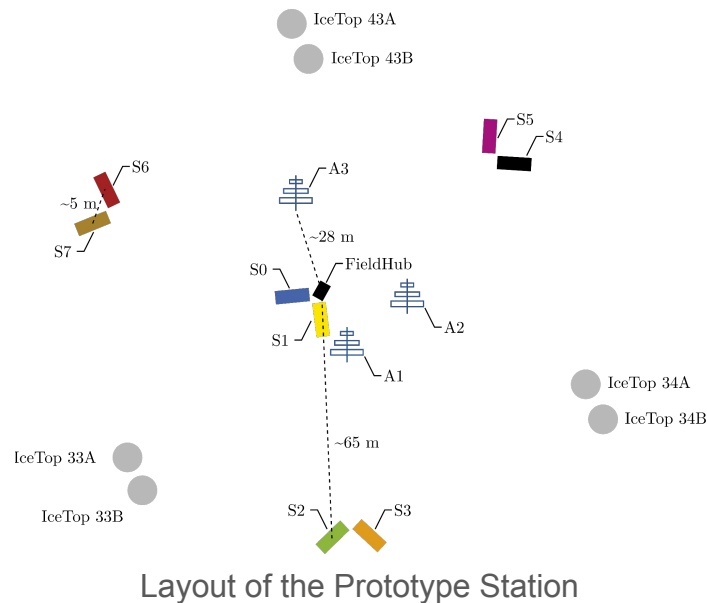
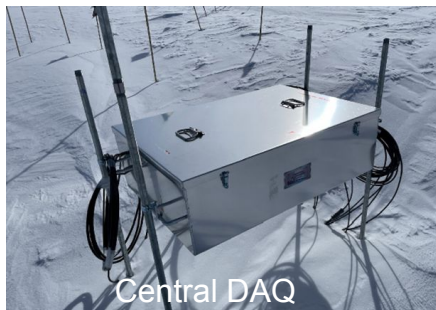
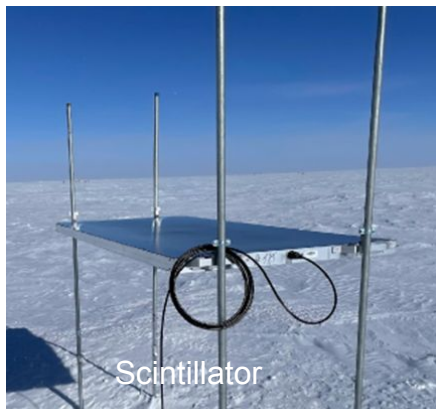


Plot credit: J. Saffer  
Snow accumulation over time on IceTop

Rising **snow accumulation** driven by winds

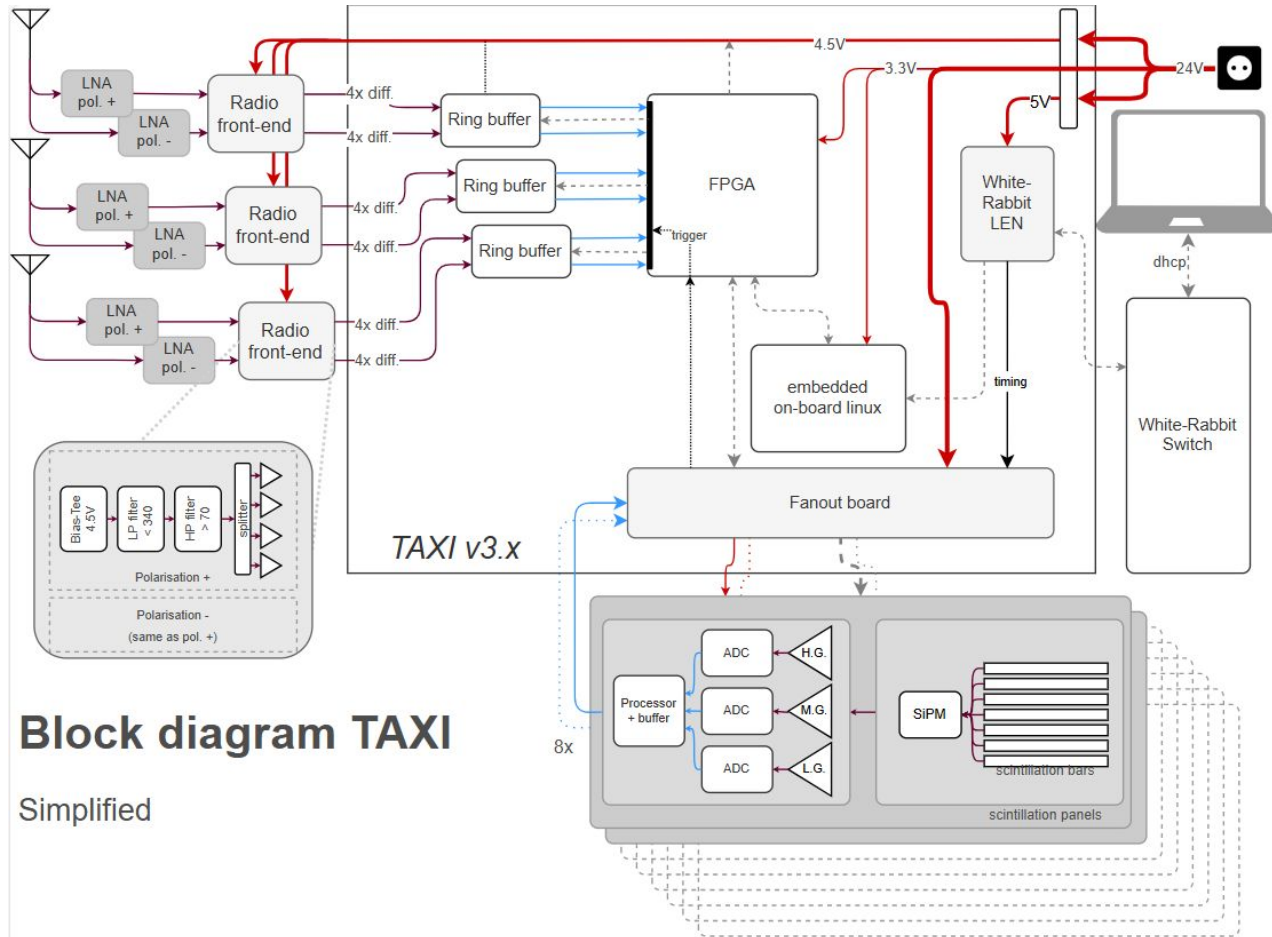
- Reduced efficiency detecting low energy cosmic rays
- Higher uncertainty in air-shower reconstruction

# The prototype station of the Surface Array Enhancement

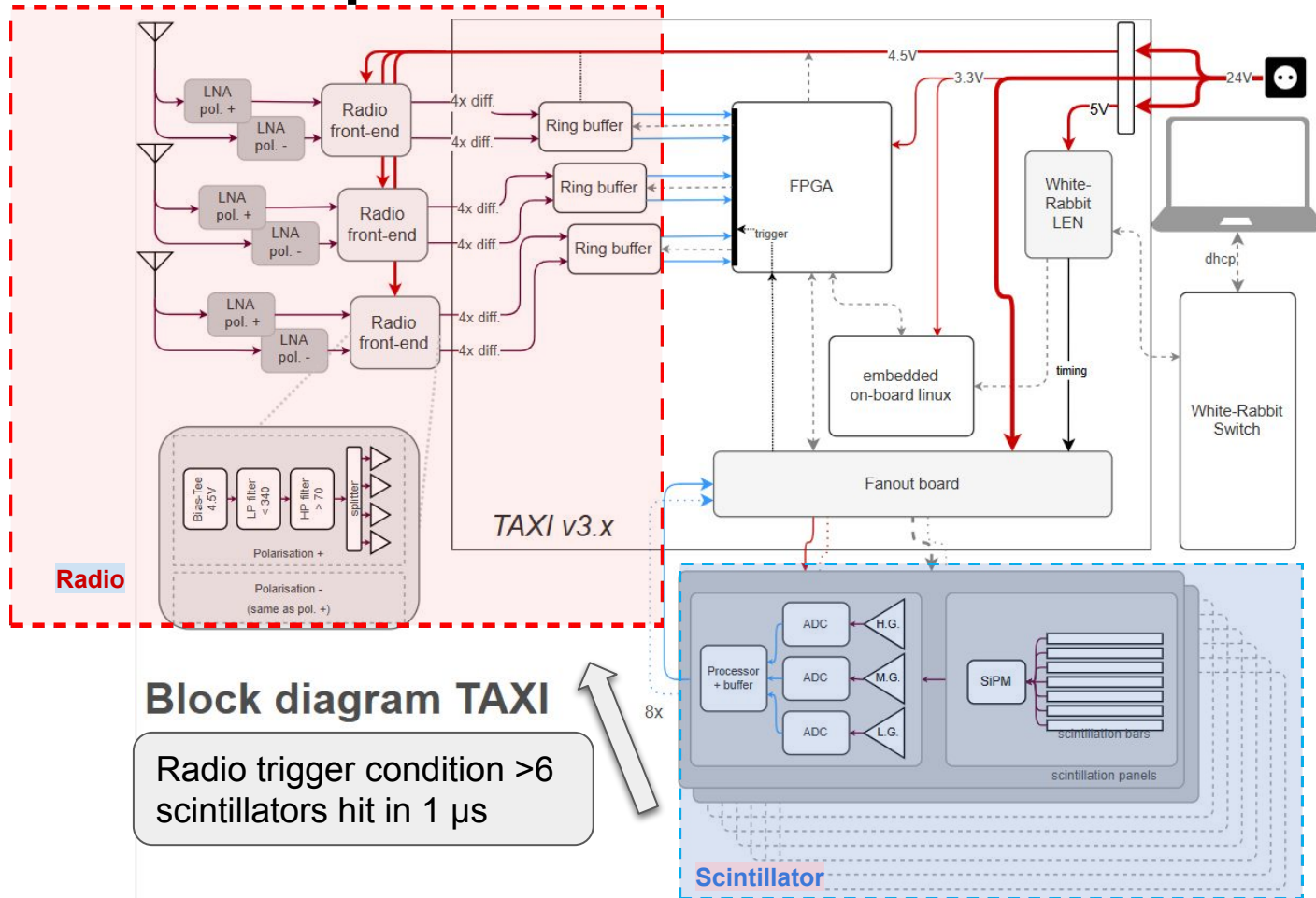


- 2 types of detectors -
  - 3 antennas and 8 scintillation detectors per station
  - Easy to elevate and cost-effective
- Radio operational band 70-350 MHz
- Deployed in 2020, upgraded in 2023

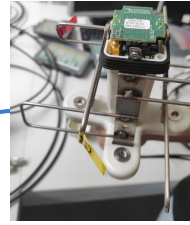
# The Data Acquisition of the SAE



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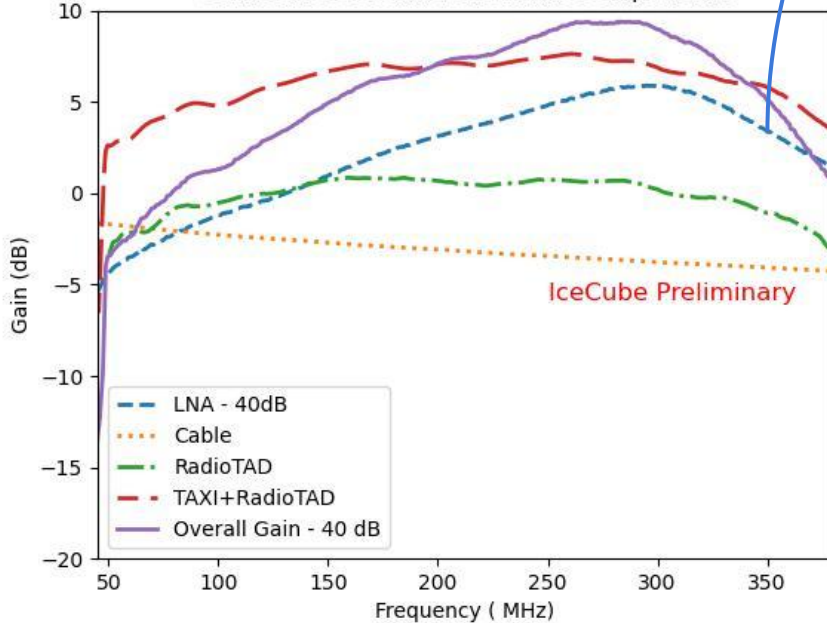


# Radio signal chain

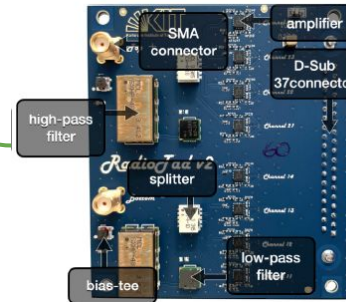
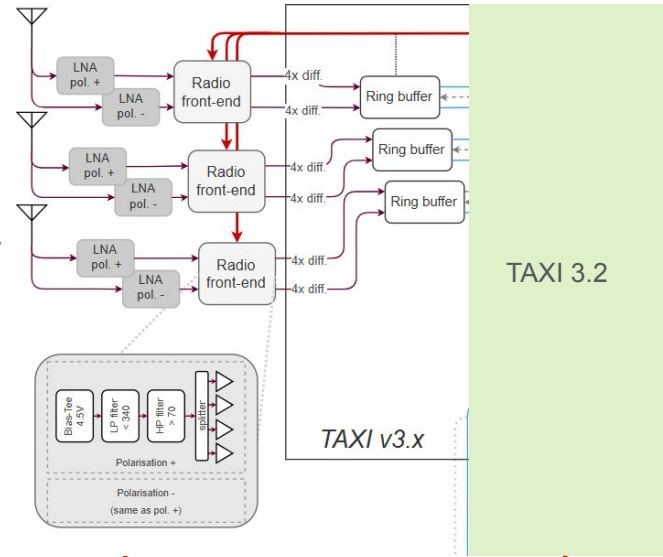


Low-Noise Amplifier

Gain curves of different Radio Components



IceCube Preliminary

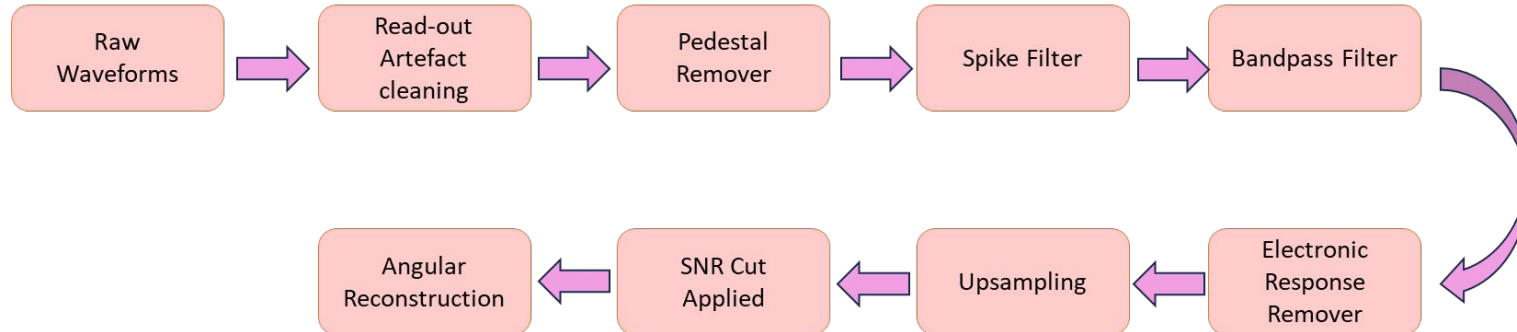


RadioTAD

# Air shower identification with the Prototype Station

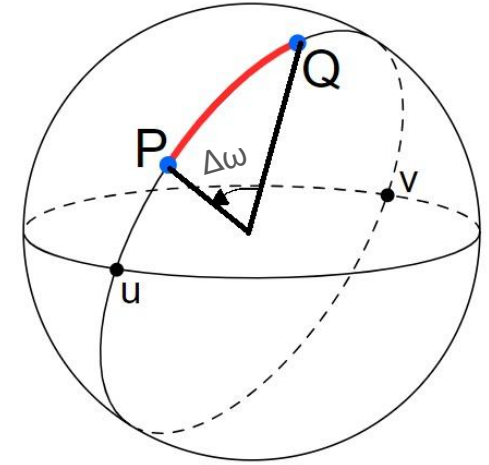
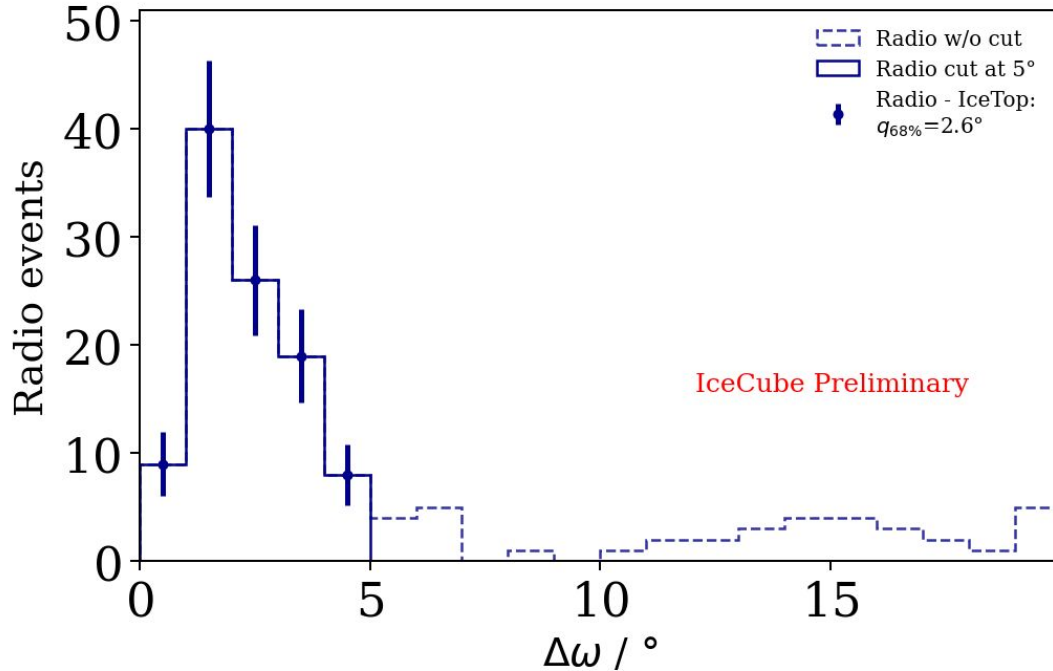
- Dataset : Jan - Jun 2022 (4 months considered) due to differences in data taking
- Radio coincidence identified with IceTop and scintillation panels - in a  $2\mu\text{s}$  window during processing.
- Cleaned for artefacts and RFI, filtered to 100-230 MHz, electronic response removed
- SNR cut computed from background data to reject 95% background per antenna
- Signal required in all 3 antennas

Background is recorded with a fixed rate trigger.  
95% of background for each antenna is rejected by checking the SNR distribution.



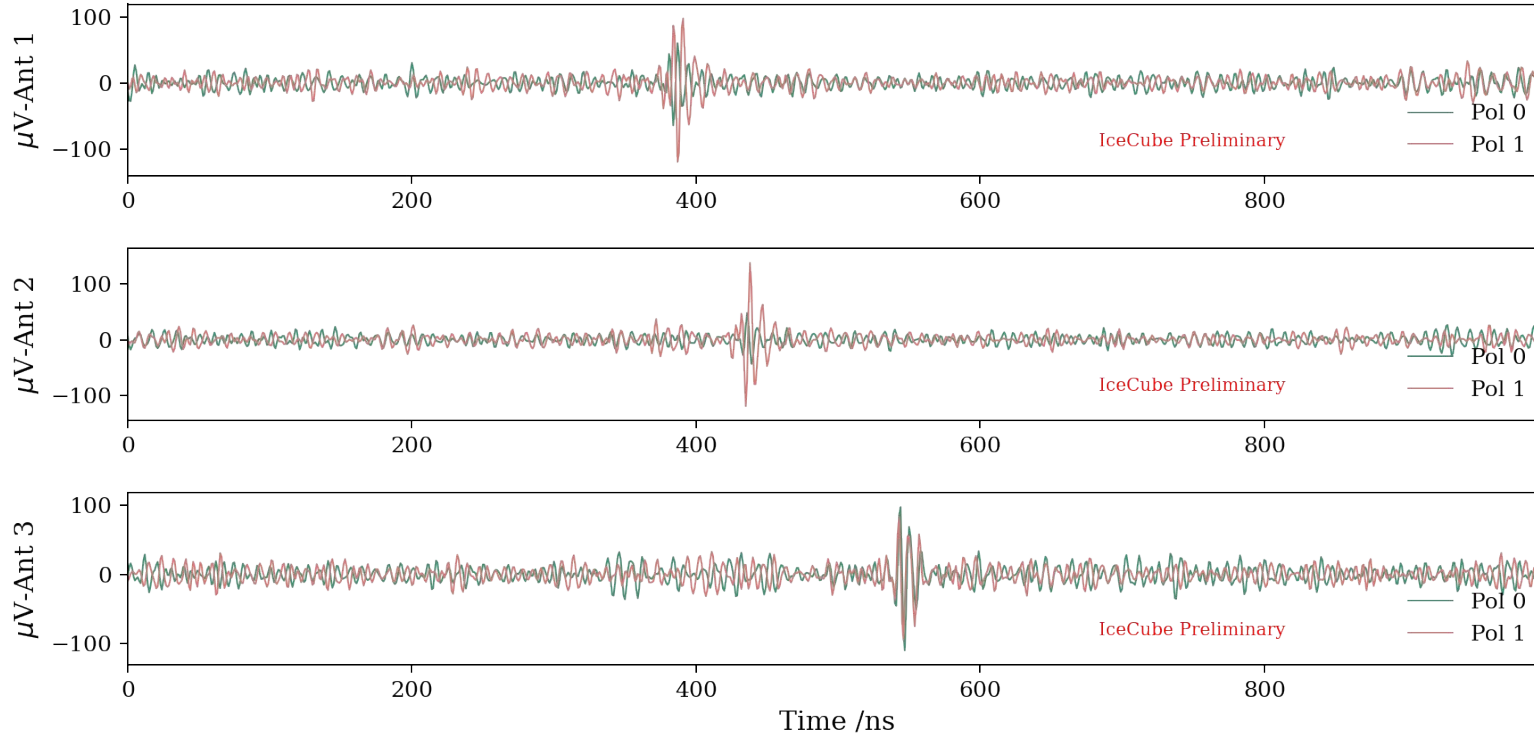


# Difference in opening angle w.r.t IceTop

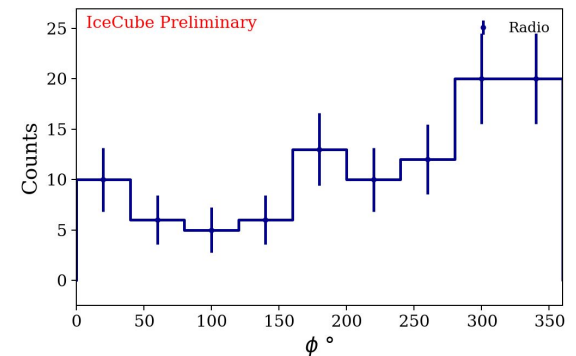
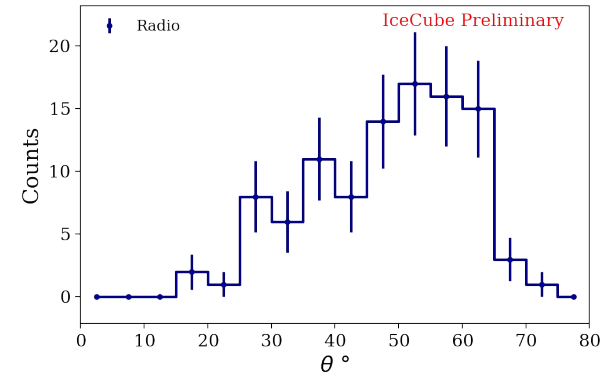
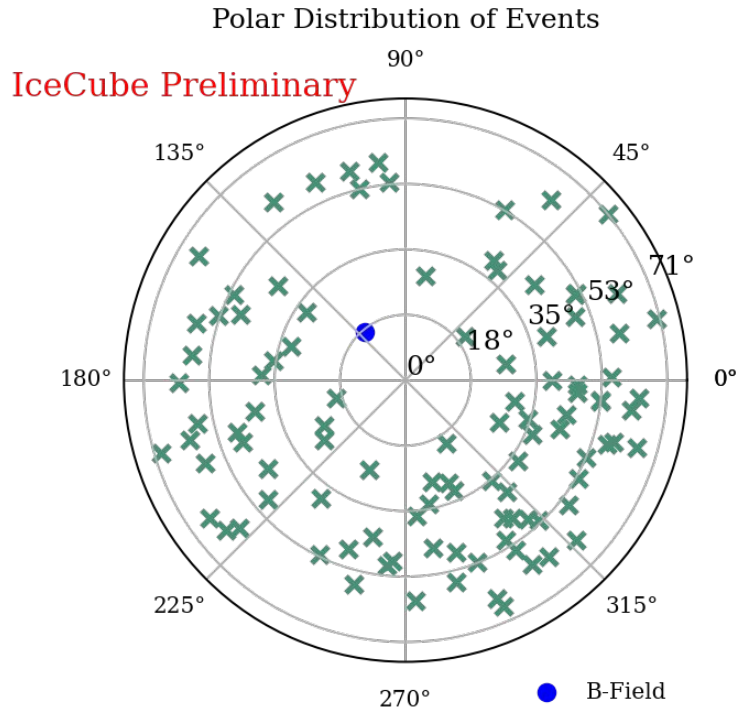


- Cut when opening angle between the direction reconstructed by IceTop and radio  $\Delta\omega \geq 5^\circ$
- Total number of events after cut with the SNR method= 104 (341 total identified events after SNR cut.)

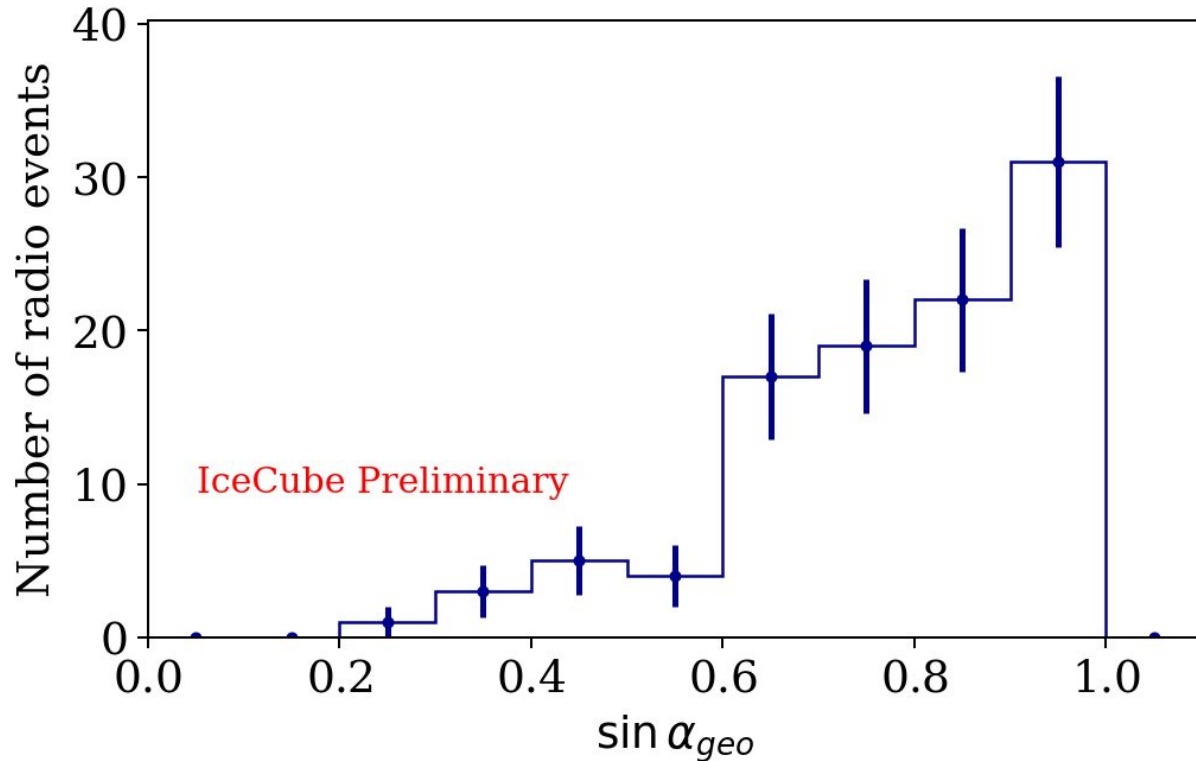
# Example air-shower event in radio



# Distribution of all radio events in the sky

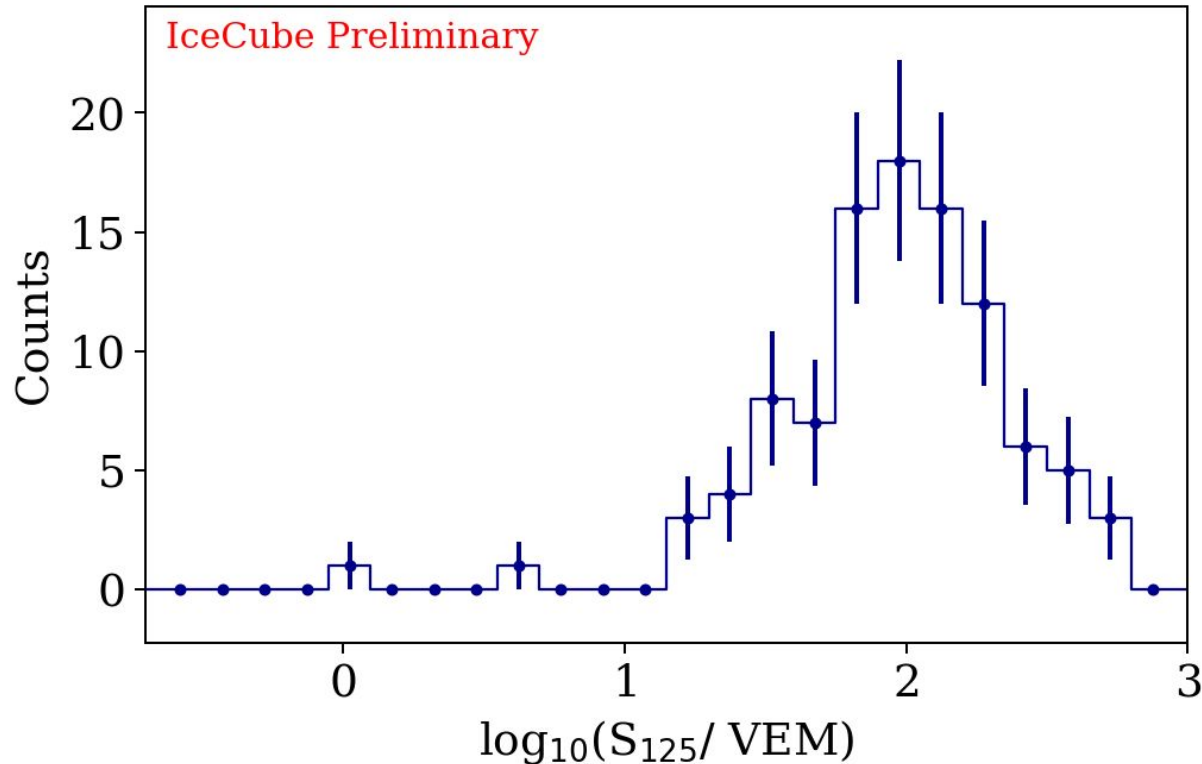


# Scaling of radio events with the geomagnetic angle



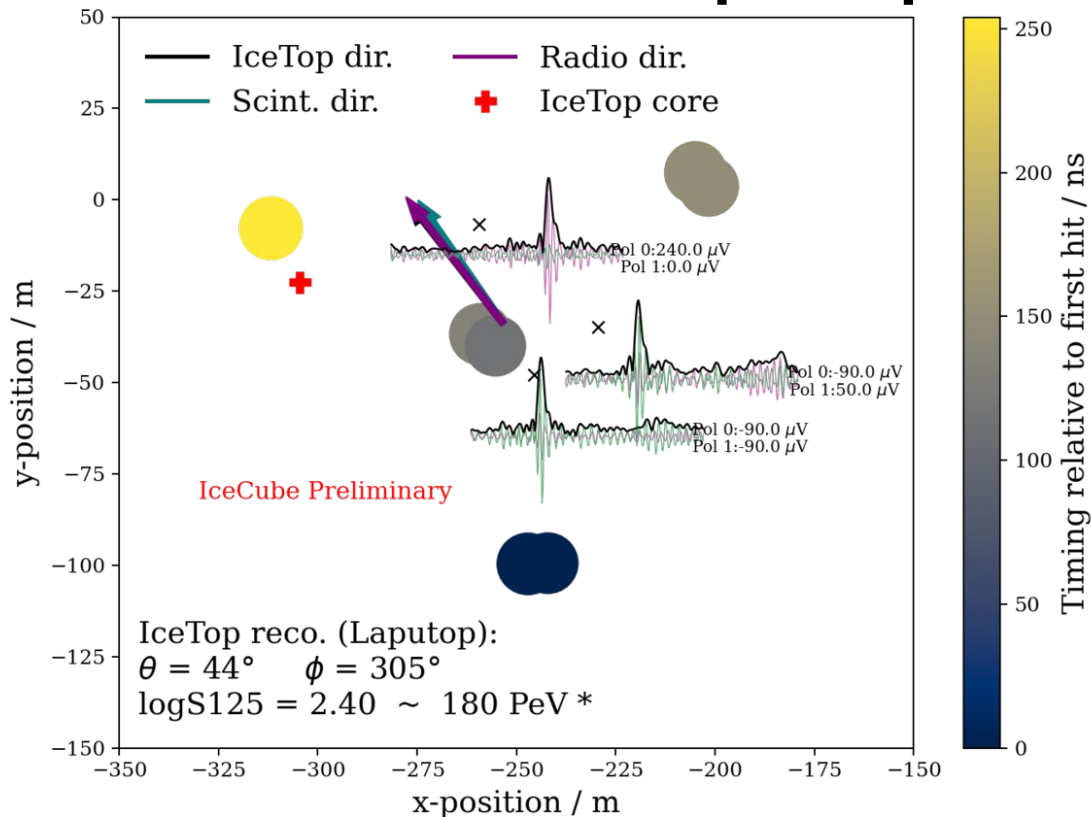
- Larger radio signal with increasing geomagnetic angle - we see more events as expected;

# S125 Distribution of all events



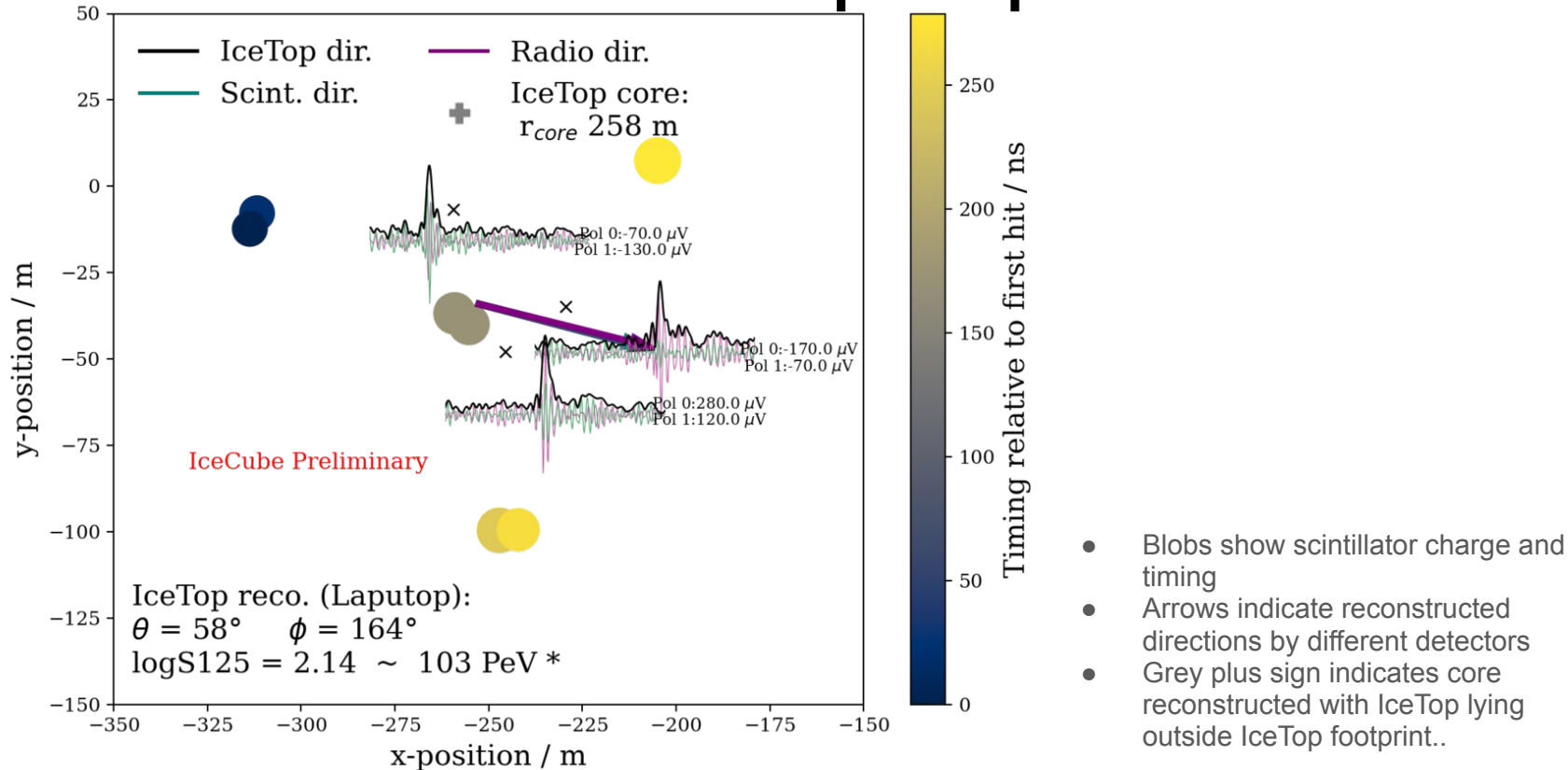
- $S_{125}$  - energy proxy for IceTop
  - Calibrated for quasi-vertical showers
  - 1 VEM  $\sim$  1PeV
- $S_{125}$  corresponds to an LDF value taken at the reference distance of 125 m from the shower axis.

# Example 3-fold coincidence events with core reconstructed in IceTop footprint



- Blobs show scintillator charge and timing
- Arrows indicate reconstructed directions by different detectors
- Red plus sign indicates core reconstructed with IceTop.

# Example 3-fold coincidence events with core reconstructed outside IceTop footprint



# Conclusion

- The antennas at the Pole are the first operating in a 70 - 350 MHz and the analysis is carried out in the 100 - 280 MHz.
- The radio antennas of the prototype station at the Pole are fully functional and detecting air showers continuously with energies starting from 10 PeV.
- We have detected over 100 events in a 4 month time period.

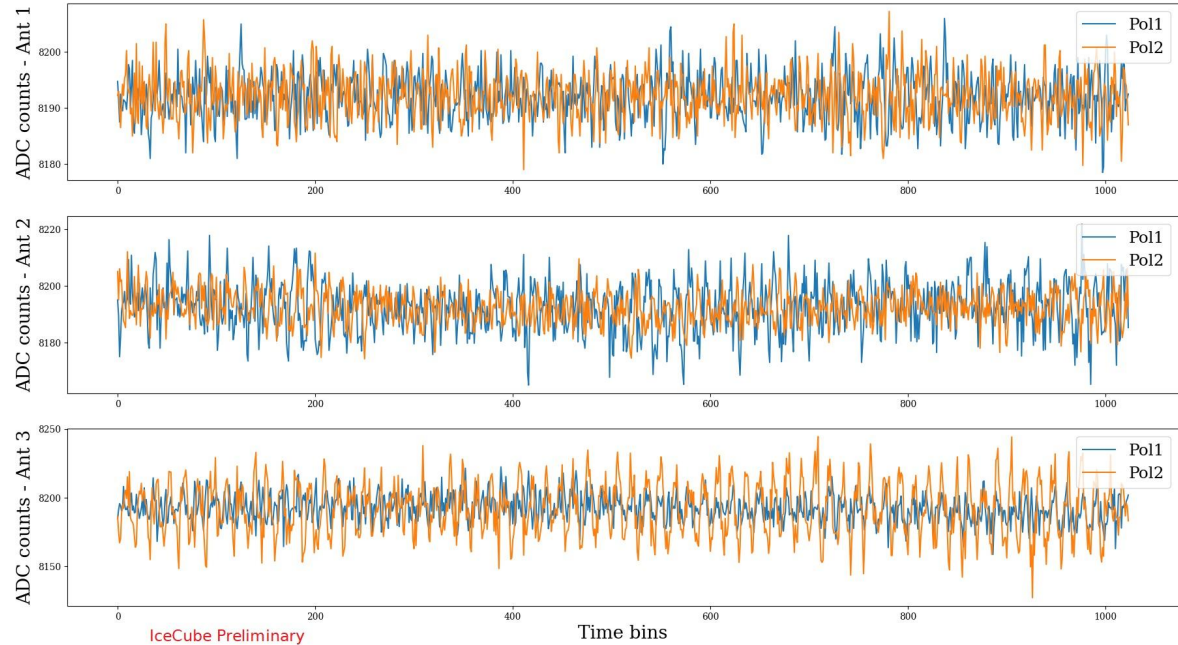
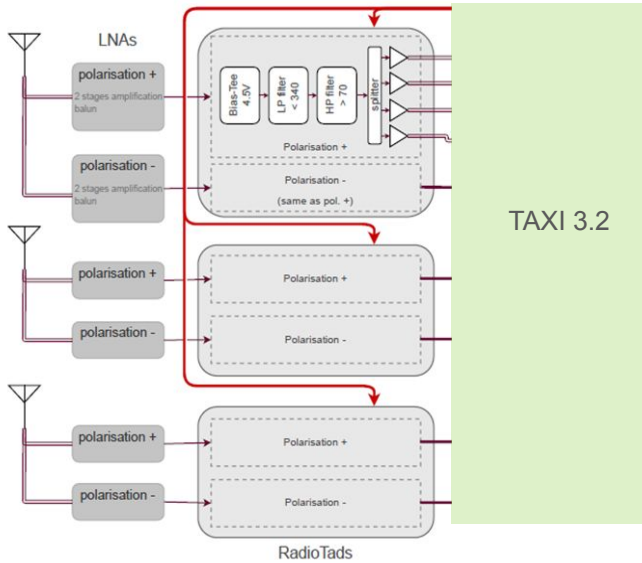
# Outlook

- The surface enhancement will be a unique detector capable of giving insights into cosmic ray science with its multi-detector systems.
- The station layout is used in the design of the IceCube-Gen2 surface array with a footprint increase by a factor of 8 over IceTop.



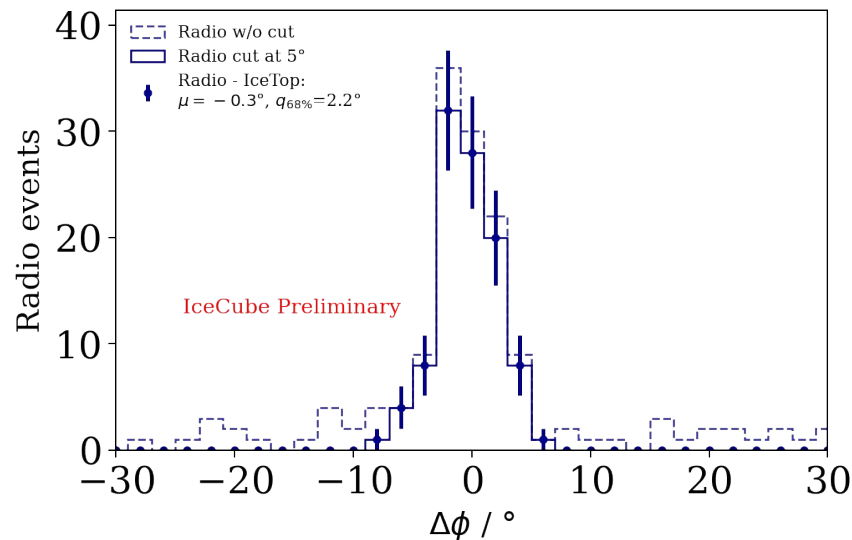
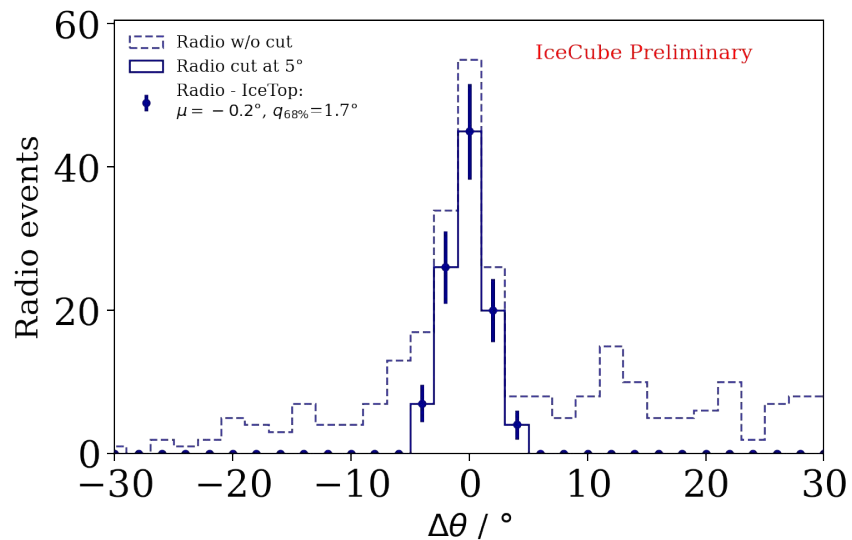
**Thank you!**

# Average background traces as measured by the DAQ

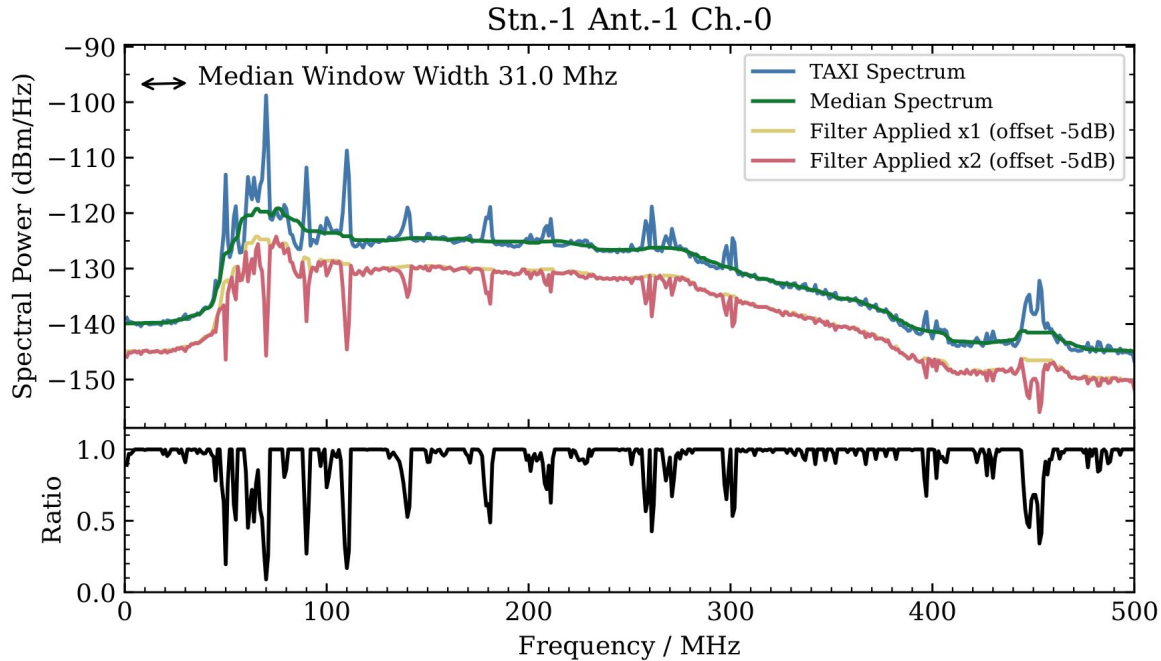


- Measured in 3 antennas over 2 polarizations
- Averaged over one day

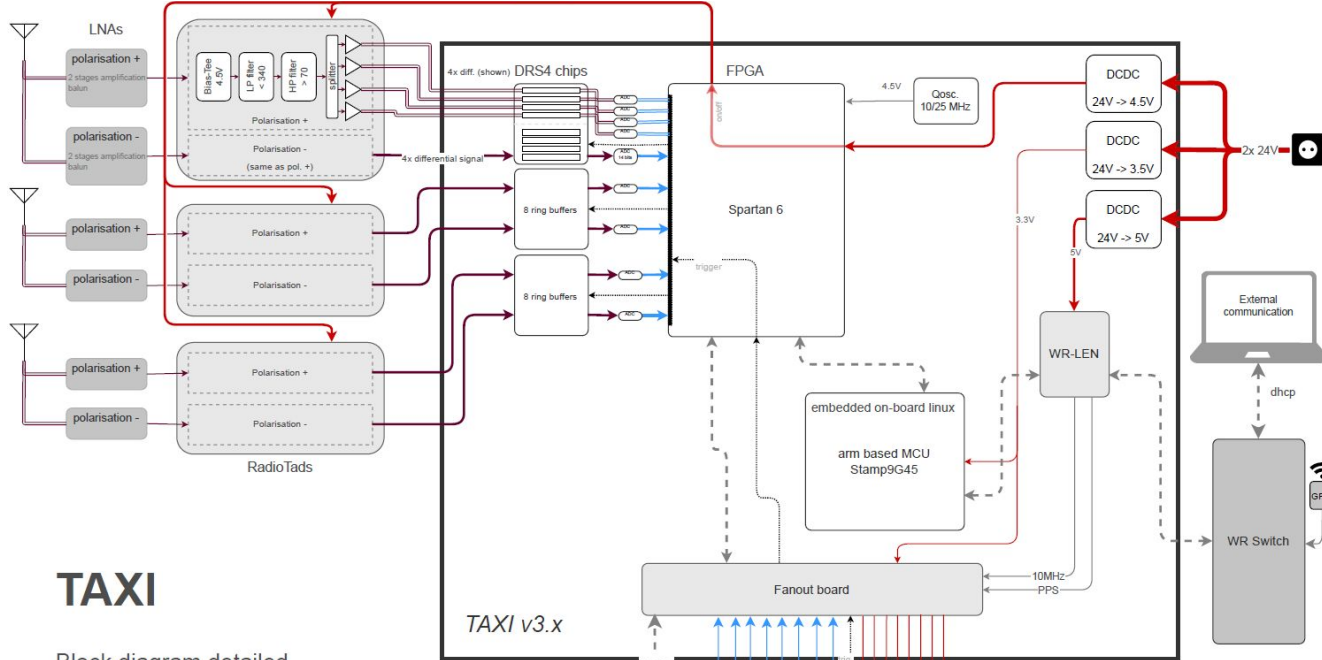
# Difference in reconstructed zenith and azimuth w.r.t IceTop



# Spike Filter for RFI cleaning



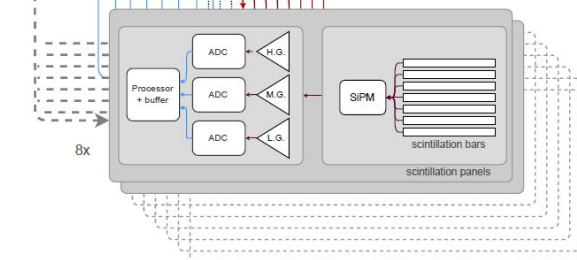
# The Data Acquisition of the SAE (detailed)



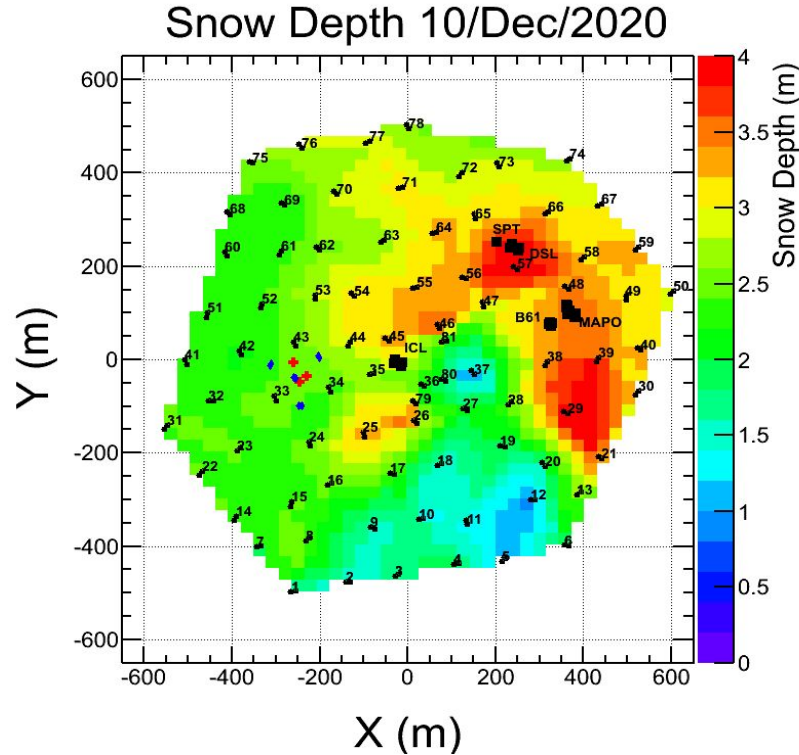
**TAXI**

Block diagram detailed

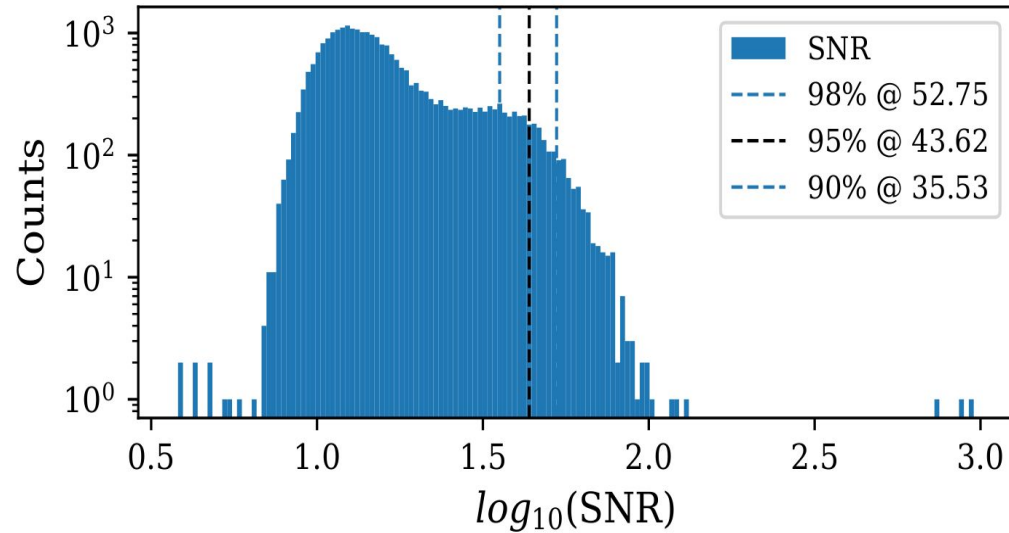
Legend :



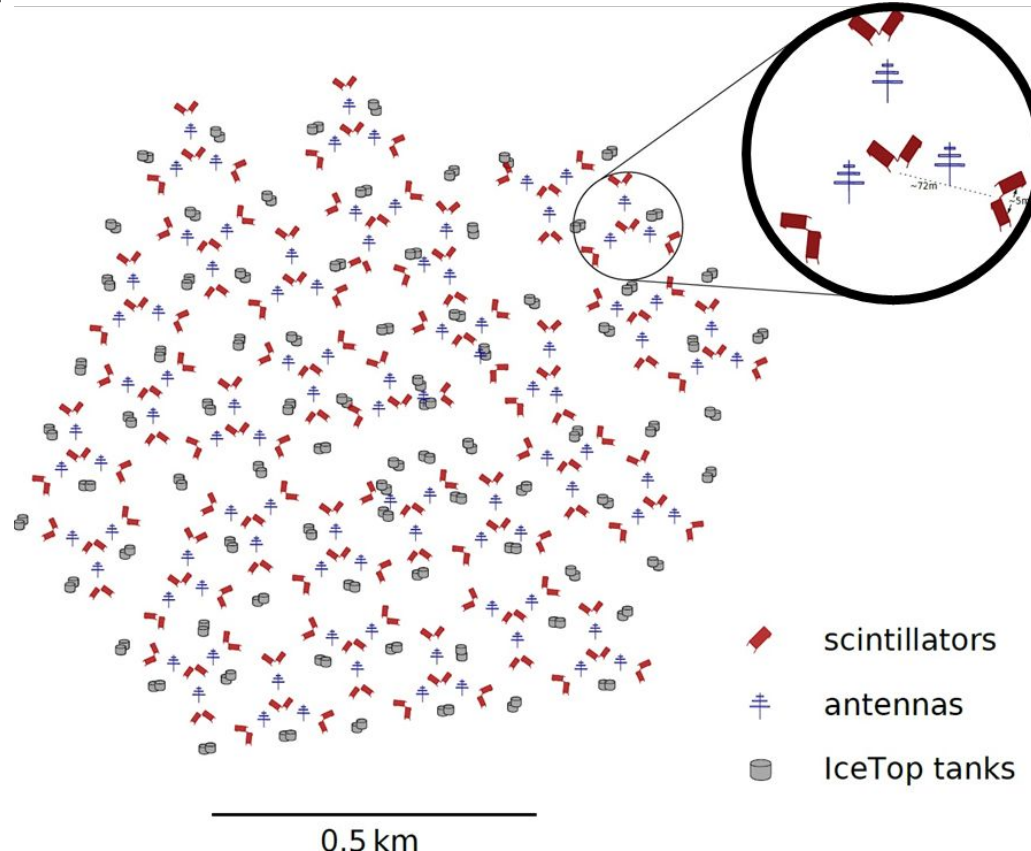
# Snow accumulation and the location of the prototype station



# SNR distribution of Antenna 1

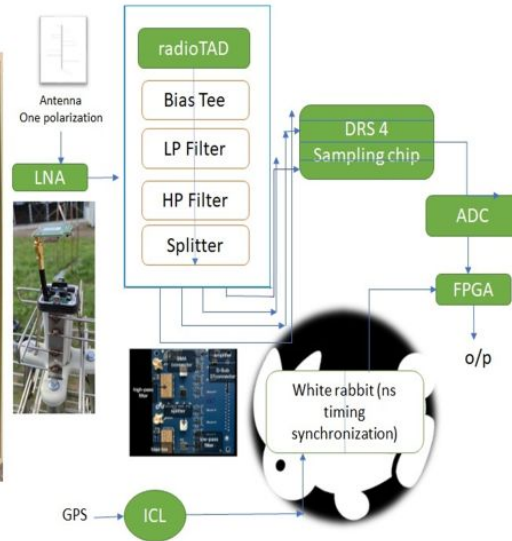


# Actual proposed layout of the enhancement and a single station

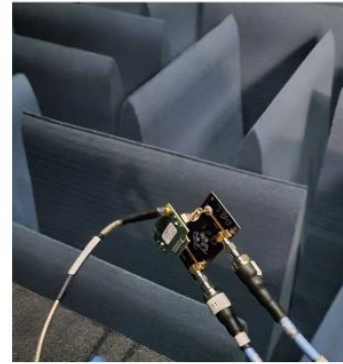




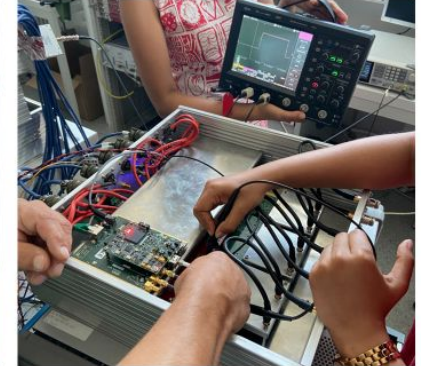
# Methods of calibration of different components



Taxi Board (Left) and Schematics of the radio signal chain in the TAXI (Right)



Calibration of LNA in an anechoic chamber  
Picture Credit: Sasha Novikov



Checking the udaq signal during timing measurements



Gain characterization for the radio signal across the TAXI board with a known input signal